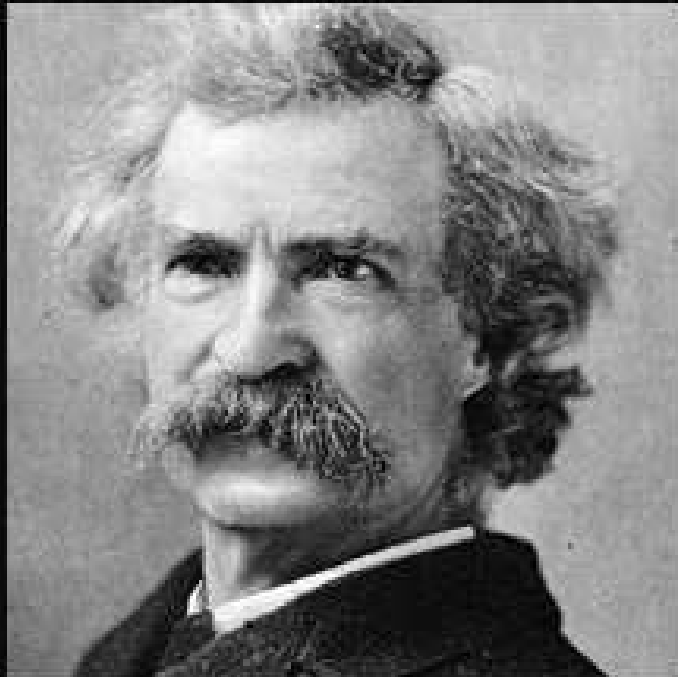


Some original SUSY literature:



The reports of my death have
been greatly exaggerated.

~ Mark Twain

Where We Will Discover SUSY

Sven Heinemeyer, IFT/IFCA (CSIC, Madrid/Santander)

Singapore, 03/2018

- Motivation & Models
- Results and predictions in GUT based models
- Results and predictions in the pMSSM11
- $\phi_{96} \rightarrow \gamma\gamma$
- Conclusions

1. Motivation & Models

Fact:

The SM cannot be the ultimate theory!

1. gravity is not included
 2. the hierarchy problem
 3. Dark Matter is not included
 4. neutrino masses are not included
 5. anomalous magnetic moment of the muon shows a $\sim 4\sigma$ discrepancy
- ⇒ Time to get ready for BSM physics

Which model should we focus on?

Some “recent” measurements:

- top quark mass
- Higgs boson mass
- Higgs boson “couplings”
- Dark Matter (properties)

Which model should we focus on?

Some “recent” measurements:

- top quark mass
- Higgs boson mass
- Higgs boson “couplings”
- Dark Matter (properties)

Simple SUSY models predicted correctly:

- top quark mass
- Higgs boson mass
- Higgs boson “couplings”
- Dark Matter (properties)

Which model should we focus on?

Some “recent” measurements:

- top quark mass
- Higgs boson mass
- Higgs boson “couplings”
- Dark Matter (properties)

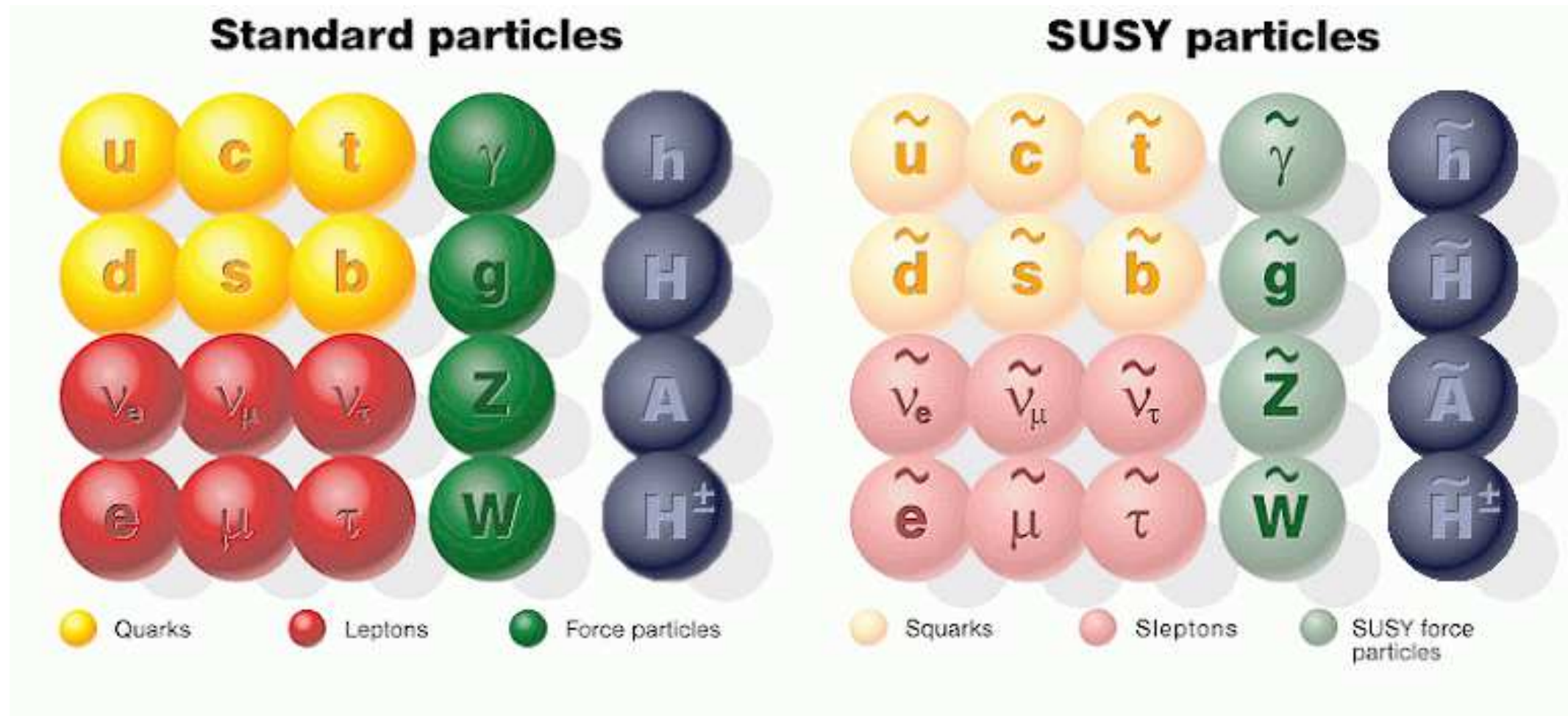
Simple SUSY models predicted correctly:

- top quark mass
- Higgs boson mass
- Higgs boson “couplings”
- Dark Matter (properties)

⇒ **good motivation to look at SUSY! :-)**

The Minimal Supersymmetric Standard Model (MSSM)

Superpartners for Standard Model particles



Problem in the MSSM: more than 100 free parameters

Nobody(?) believes that a model describing nature has so many free parameters!

Enlarged Higgs sector: Two Higgs doublets

$$H_1 = \begin{pmatrix} H_1^1 \\ H_1^2 \end{pmatrix} = \begin{pmatrix} v_1 + (\phi_1 + i\chi_1)/\sqrt{2} \\ \phi_1^- \end{pmatrix}$$
$$H_2 = \begin{pmatrix} H_2^1 \\ H_2^2 \end{pmatrix} = \begin{pmatrix} \phi_2^+ \\ v_2 + (\phi_2 + i\chi_2)/\sqrt{2} \end{pmatrix}$$

$$V = m_1^2 H_1 \bar{H}_1 + m_2^2 H_2 \bar{H}_2 - m_{12}^2 (\epsilon_{ab} H_1^a H_2^b + \text{h.c.})$$
$$+ \underbrace{\frac{g'^2 + g^2}{8}}_{\text{gauge couplings, in contrast to SM}} (H_1 \bar{H}_1 - H_2 \bar{H}_2)^2 + \underbrace{\frac{g^2}{2}}_{\text{gauge couplings, in contrast to SM}} |H_1 \bar{H}_2|^2$$

physical states: h^0, H^0, A^0, H^\pm

Goldstone bosons: G^0, G^\pm

Input parameters: (to be determined experimentally)

$$\tan \beta = \frac{v_2}{v_1}, \quad M_A^2 = -m_{12}^2 (\tan \beta + \cot \beta)$$

A. Unconstrained models (MSSM):

agnostic about how SUSY breaking is achieved

no particular SUSY breaking mechanism assumed, parameterization of possible soft SUSY-breaking terms

most general case:

⇒ 105 new parameters: masses, mixing angles, phases

⇒ no model missed (within the MSSM)

⇒ $\mathcal{O}(100)$ parameters difficult to handle

B. Constrained models:

CMSSM, NUHM1, NUHM2, SU(5), mAMSB,:

assumption on the scenario that achieves spontaneous SUSY breaking

⇒ prediction for soft SUSY-breaking terms
in terms of small set of parameters

⇒ easy to handle

GUT based models: 1.) CMSSM (sometimes wrongly called mSUGRA):

⇒ Scenario characterized by

$$m_0, m_{1/2}, A_0, \tan \beta, \text{sign } \mu$$

m_0 : universal scalar mass parameter
 $m_{1/2}$: universal gaugino mass parameter
 A_0 : universal trilinear coupling
 $\tan \beta$: ratio of Higgs vacuum expectation values
 $\text{sign}(\mu)$: sign of supersymmetric Higgs parameter

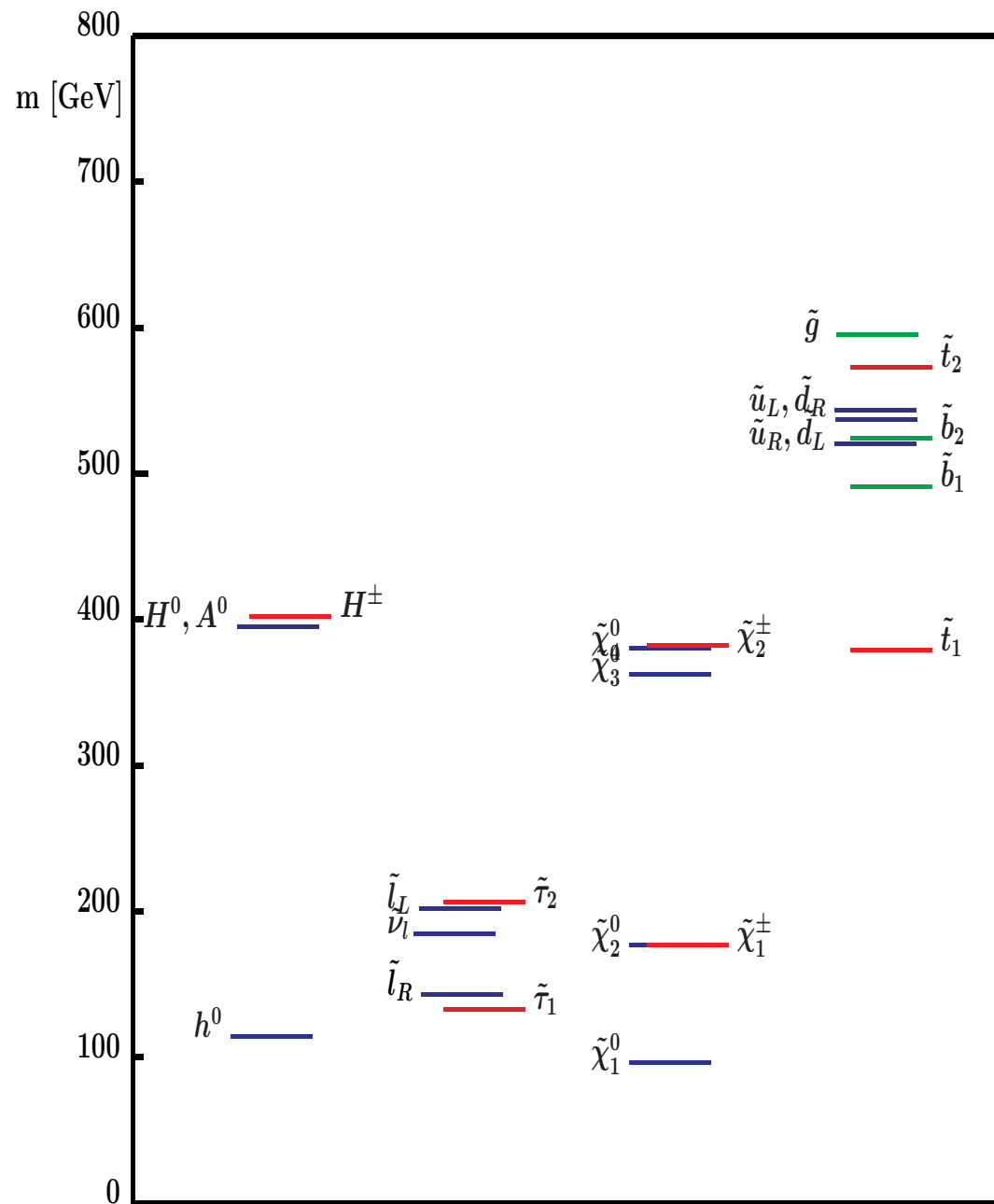
} at the GUT scale

⇒ particle spectra from renormalization group running to weak scale

⇒ Lightest SUSY particle (LSP) is the lightest neutralino ⇒ DM!

“Typical” CMSSM scenario
(SPS 1a benchmark scenario):

Close connection between
all the sectors



GUT based models: 2.) NUHM1: (Non-universal Higgs mass model)

Assumption: no unification of scalar fermion and scalar Higgs parameter at the GUT scale

⇒ effectively M_A as free parameters at the EW scale

⇒ Scenario characterized by

$$m_0, m_{1/2}, A_0, \tan \beta, \text{sign } \mu \text{ and } M_A$$

GUT based models: 3.) NUHM2: (Non-universal Higgs mass model 2)

Assumption: no unification of scalar Higgs parameter at the GUT scale

⇒ effectively M_A and μ as free parameters at the EW scale

⇒ Scenario characterized by

$$m_0, m_{1/2}, A_0, \tan \beta, \mu \text{ and } M_A$$

GUT based models: 4.) SU(5) GUT:

Assumption I:

no unification of scalar Higgs parameter at the GUT scale

(\Rightarrow effectively M_A and μ as free parameters at the EW scale)

Assumption II:

$$(q_L, u_L^c, e_L^c)_i \in \mathbf{10}_i, \quad (\ell_L, d_L^c)_i \in \bar{\mathbf{5}}_i$$

\Rightarrow Scenario characterized by

$$m_5, m_{10}, m_{1/2}, A_0, \tan \beta, m_{H_u}, m_{H_d}$$

GUT based models: 5.) mAMSB:

mAMSB scenario characterized by

$$m_{3/2}, m_0, \tan \beta, \text{sign}(\mu)$$

$m_{3/2} = \langle F \rangle / M_{\text{Planck}}$: overall scale of SUSY particle masses

m_0 : phenomenological parameter: universal scalar mass term introduced in order to keep squares of slepton masses positive

typical feature: very small neutralino–chargino mass difference
 $\Rightarrow \tilde{\chi}_1^\pm \rightarrow \tilde{\chi}_1^0 + \pi^\pm$ with very soft pions

Problem: We cannot be sure about the SUSY-breaking mechanism

- ⇒ it is possible that with the CMSSM, NUHM1, NUHM2, SU(5), mAMSB we missed the “correct” mechanism
- ⇒ hint: strong connection between colored and uncolored sector
tension between low-energy EW effects and (colored) LHC searches

Problem: We cannot be sure about the SUSY-breaking mechanism

- ⇒ it is possible that with the CMSSM, NUHM1, NUHM2, SU(5), mAMSB we missed the “correct” mechanism
- ⇒ hint: strong connection between colored and uncolored sector
tension between low-energy EW effects and (colored) LHC searches

Solution: investigate also the “general MSSM”

⇒ 11 parameters are manageable ⇒ pMSSM11

- squark mass parameters: $m_{\tilde{q}_{1,2}} =: m_{\tilde{q}}, m_{\tilde{q}_3}$
- slepton mass parameter(s): $m_{\tilde{l}}, m_{\tilde{\tau}}$
- gaugino masses: M_1, M_2, M_3
- trilinear coupling: A
- Higgs sector parameters: $M_A, \tan \beta$
- Higgs mixing parameter: μ

⇒ collaborative effort of theorists and experimentalists

[Bagnaschi, Borsato, Buchmüller, Cavanaugh, Chobanova, Citron, Costa, De Roeck, Dolan, Ellis, Flücher, SH, Isidori, Liu, Lucio, Martinez Santos, Olive, Richards, Sakurai, Weiglein]

Über-code for the combination of different tools:

- Über-code original in Fortran, now re-written in C++
- tools are included as **subroutines**
- **compatibility** ensured by collaboration of authors of “MasterCode” and authors of “sub tools” **/SLHA(2)**
- sub-codes in Fortran or C++

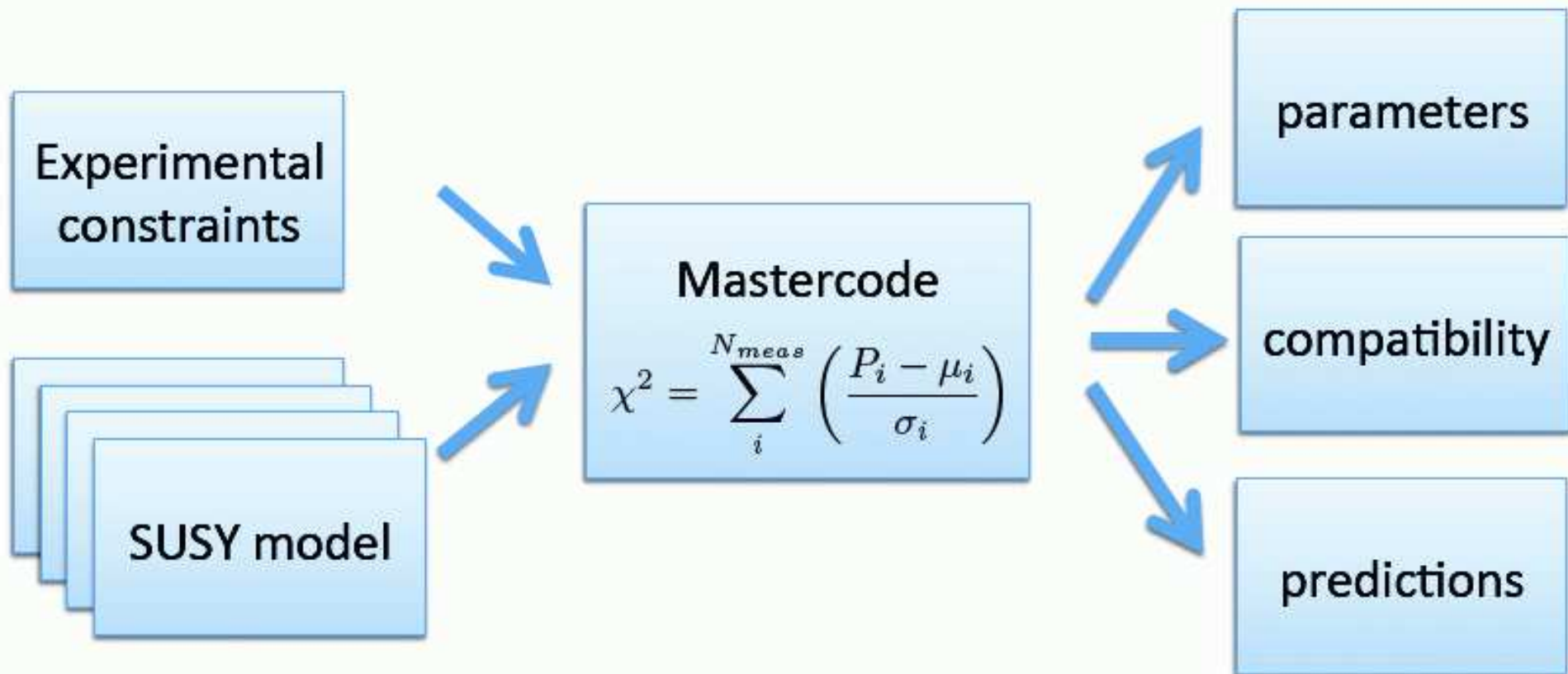
⇒ evaluate observables of one parameter point consistently with various tools

cern.ch/mastercode

The χ^2 evaluation:



Global fits of SUSY



Data we have:

- Higgs boson mass (LHC) \Rightarrow FeynHiggs

Data we have:

- Higgs boson mass (LHC) \Rightarrow FeynHiggs
- Higgs boson signal strengths (LHC) \Rightarrow HiggsSignals

Data we have:

- Higgs boson mass (LHC) \Rightarrow FeynHiggs
- Higgs boson signal strengths (LHC) \Rightarrow HiggsSignals
- Higgs boson exclusion bounds (LHC, Tevatron, LEP) \Rightarrow HiggsBounds

Data we have:

- Higgs boson mass (LHC) \Rightarrow FeynHiggs
- Higgs boson signal strengths (LHC) \Rightarrow HiggsSignals
- Higgs boson exclusion bounds (LHC, Tevatron, LEP) \Rightarrow HiggsBounds
- SUSY searches (LHC) \Rightarrow own re-cast (Fastlim approach)

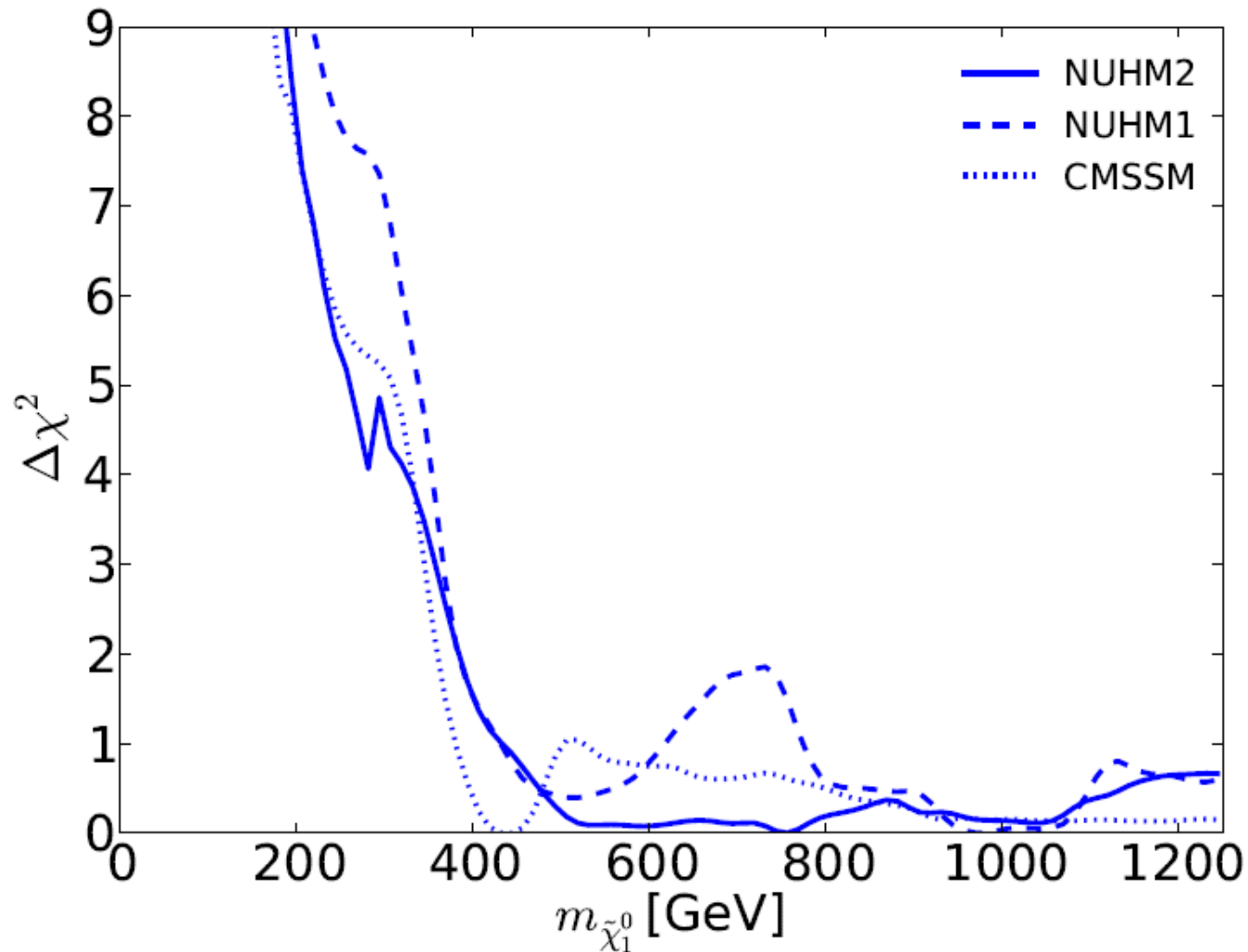
Data we have:

- Higgs boson mass (LHC) \Rightarrow FeynHiggs
- Higgs boson signal strengths (LHC) \Rightarrow HiggsSignals
- Higgs boson exclusion bounds (LHC, Tevatron, LEP) \Rightarrow HiggsBounds
- SUSY searches (LHC) \Rightarrow own re-cast (Fastlim approach)
- electroweak precision data \Rightarrow FeynWZ, FeynHiggs
- flavor data \Rightarrow SuperIso, SuFla
- astrophysical data (DM properties) \Rightarrow MicrOMEGAs, SSARD

2. Results and predictions in GUT based models

[2014]

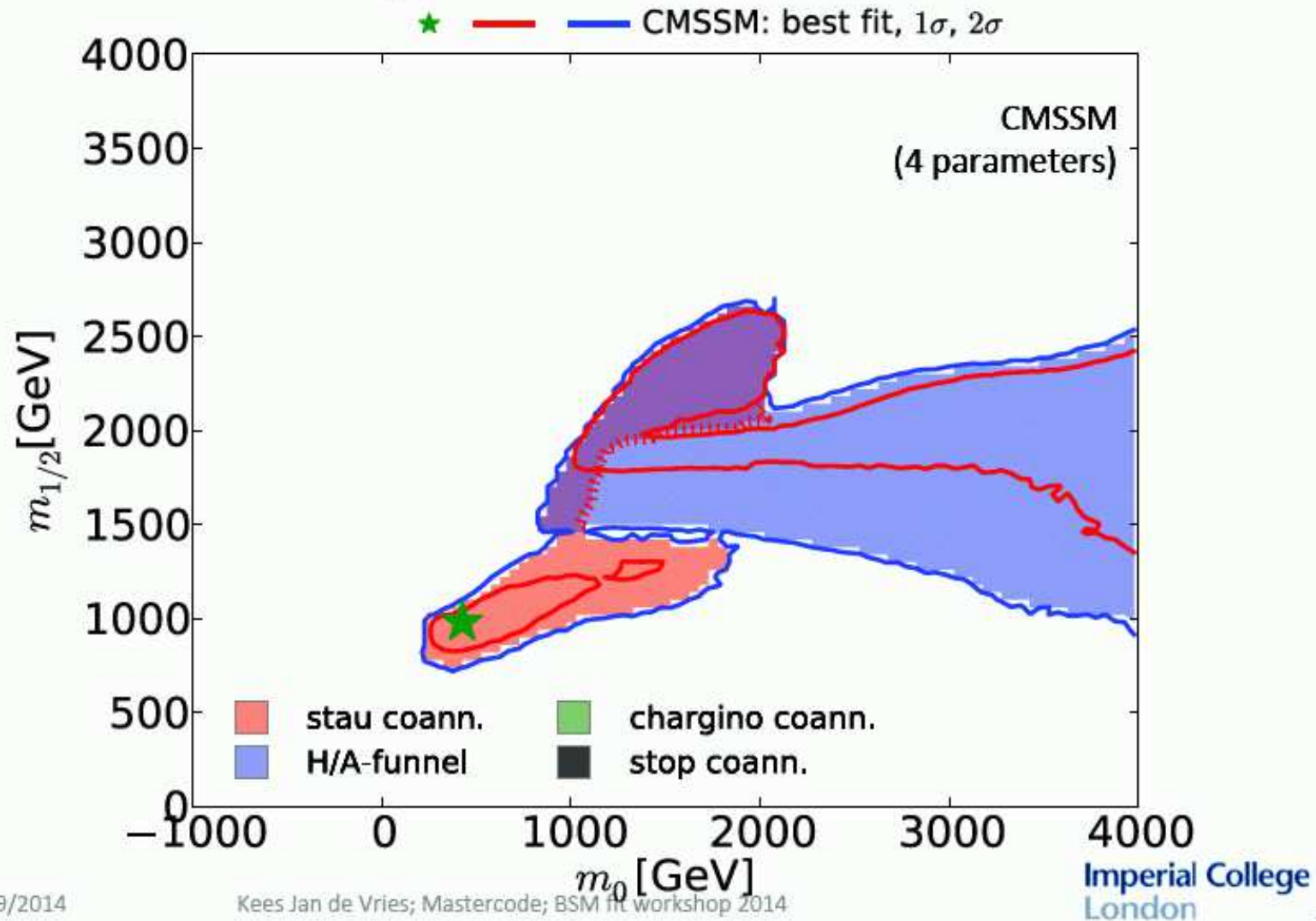
Results in the CMSSM, NUHM1, NUHM2



⇒ only very large values are favored



Mechanisms for relic dark matter density fulfillment in the CMSSM

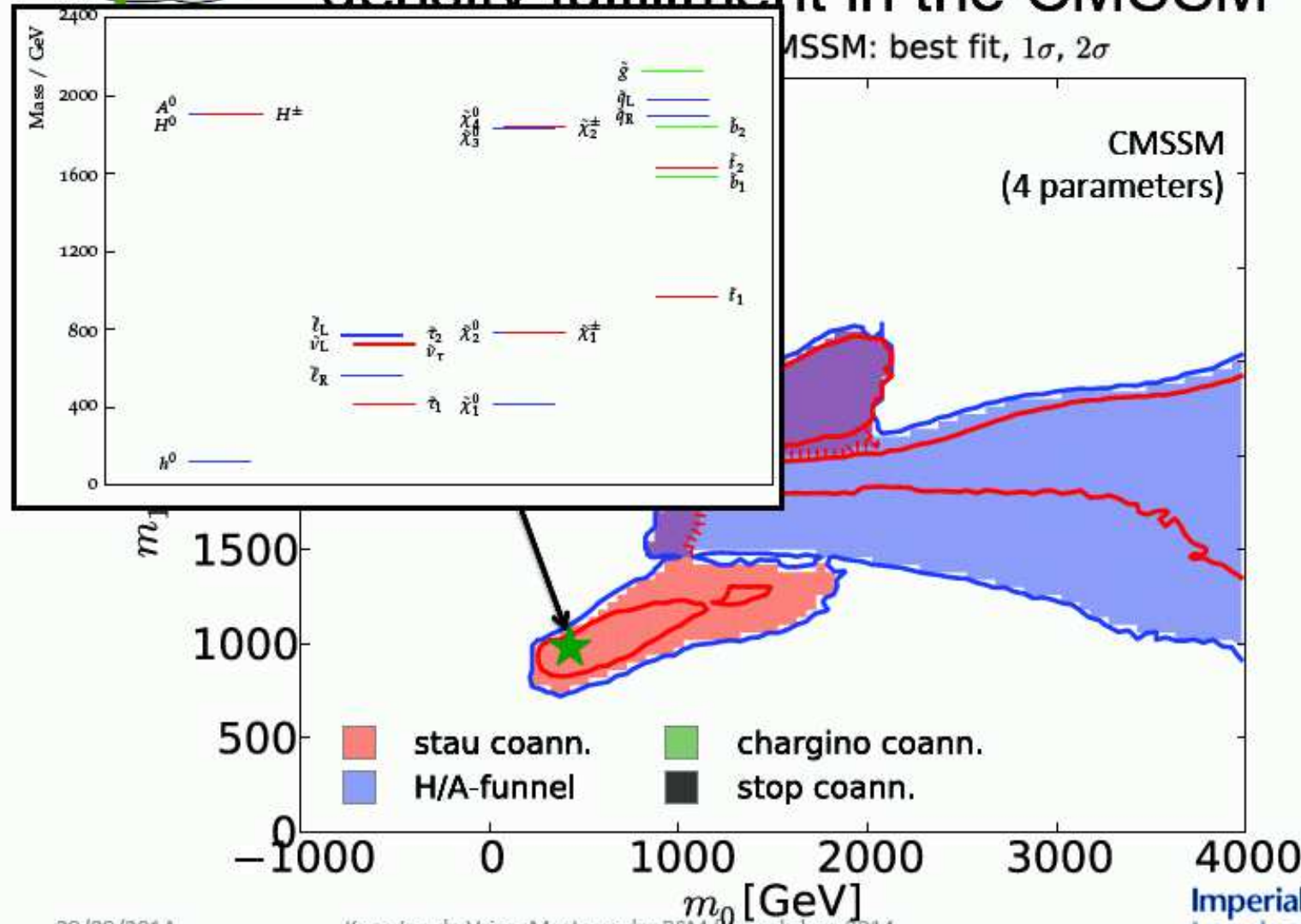


29/09/2014

Kees Jan de Vries; Mastercode; BSM fit workshop 2014



Mechanisms for relic dark matter density fulfillment in the CMSSM



29/09/2014

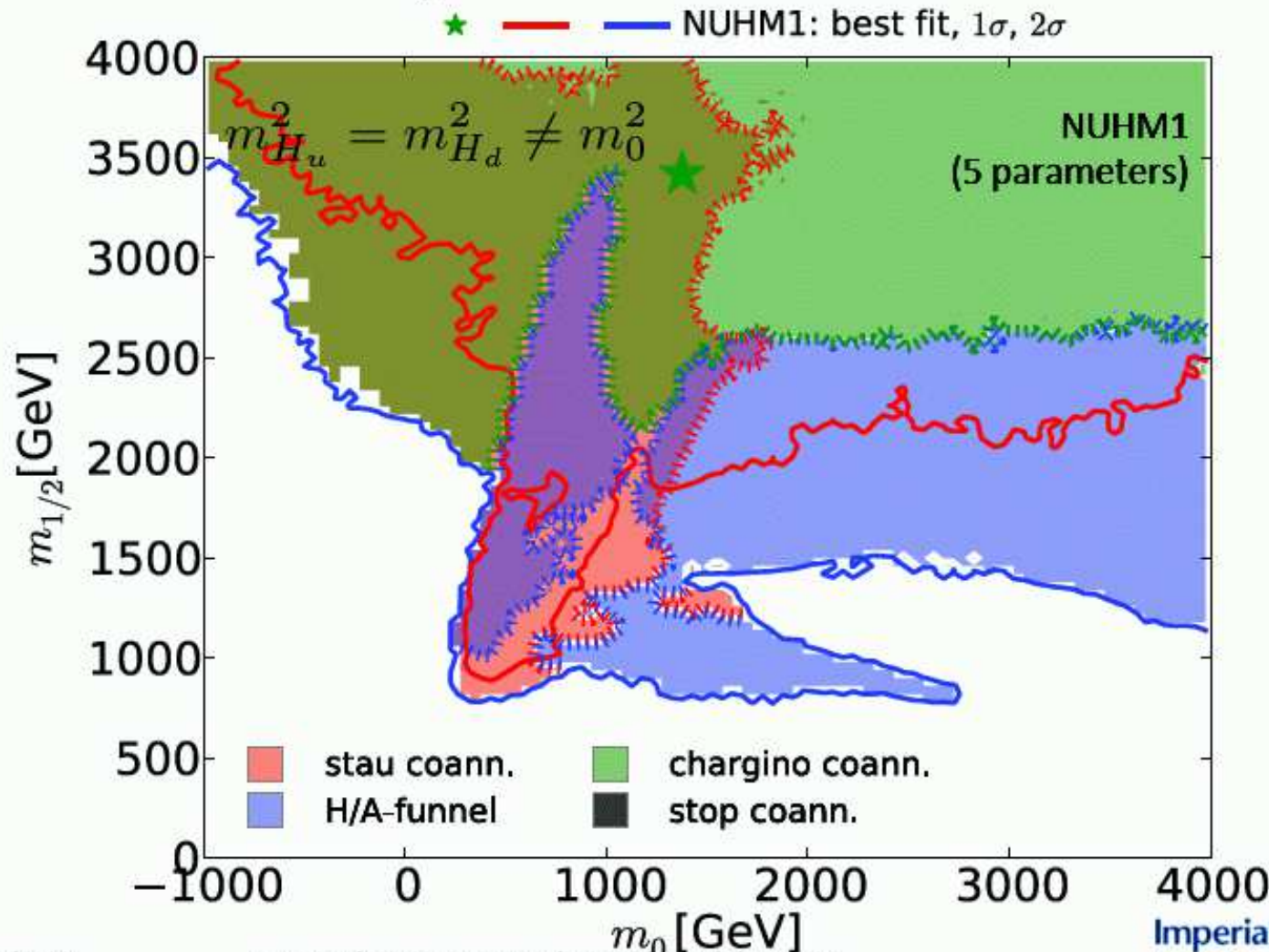
Kees Jan de Vries; Mastercode; BSM fit workshop 2014

Imperial College
London

8



Mechanisms for relic dark matter density fulfillment in the NUHM1



29/09/2014

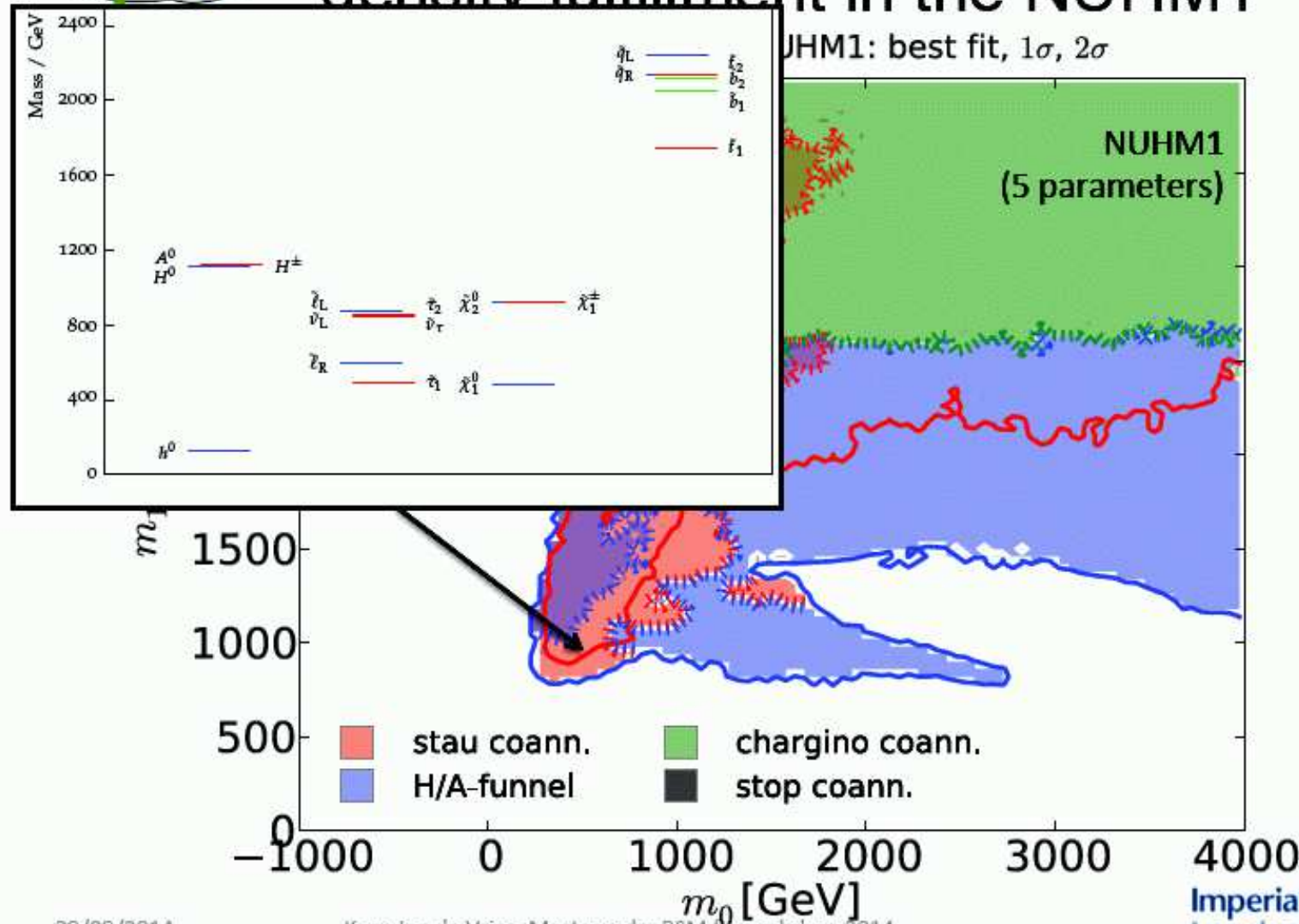
Kees Jan de Vries; Mastercode; BSM fit workshop 2014

Imperial College
London

10



Mechanisms for relic dark matter density fulfillment in the NUHM1



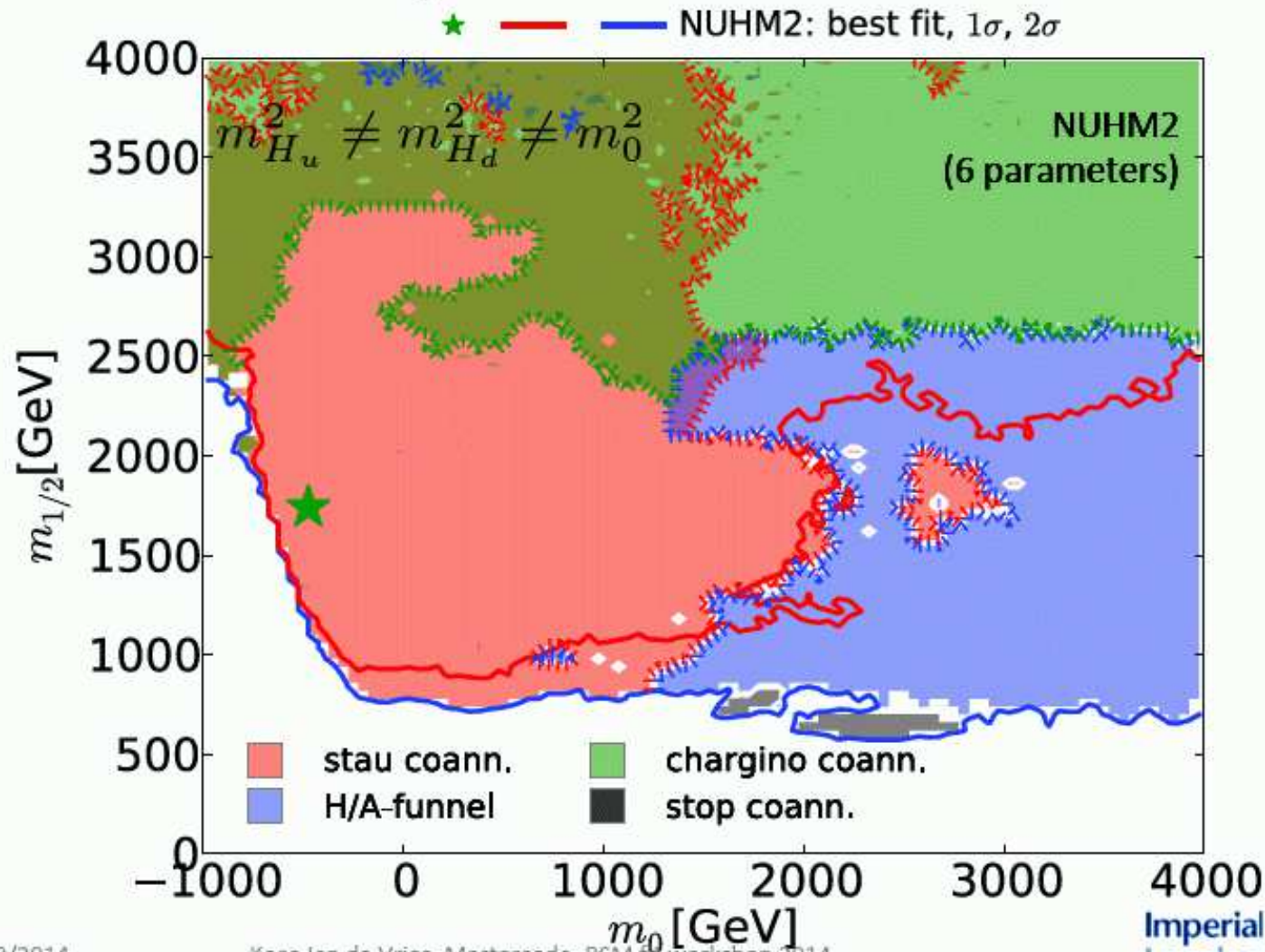
29/09/2014

Kees Jan de Vries; Mastercode; BSM fit workshop 2014

11



Mechanisms for relic dark matter density fulfillment in the NUHM2

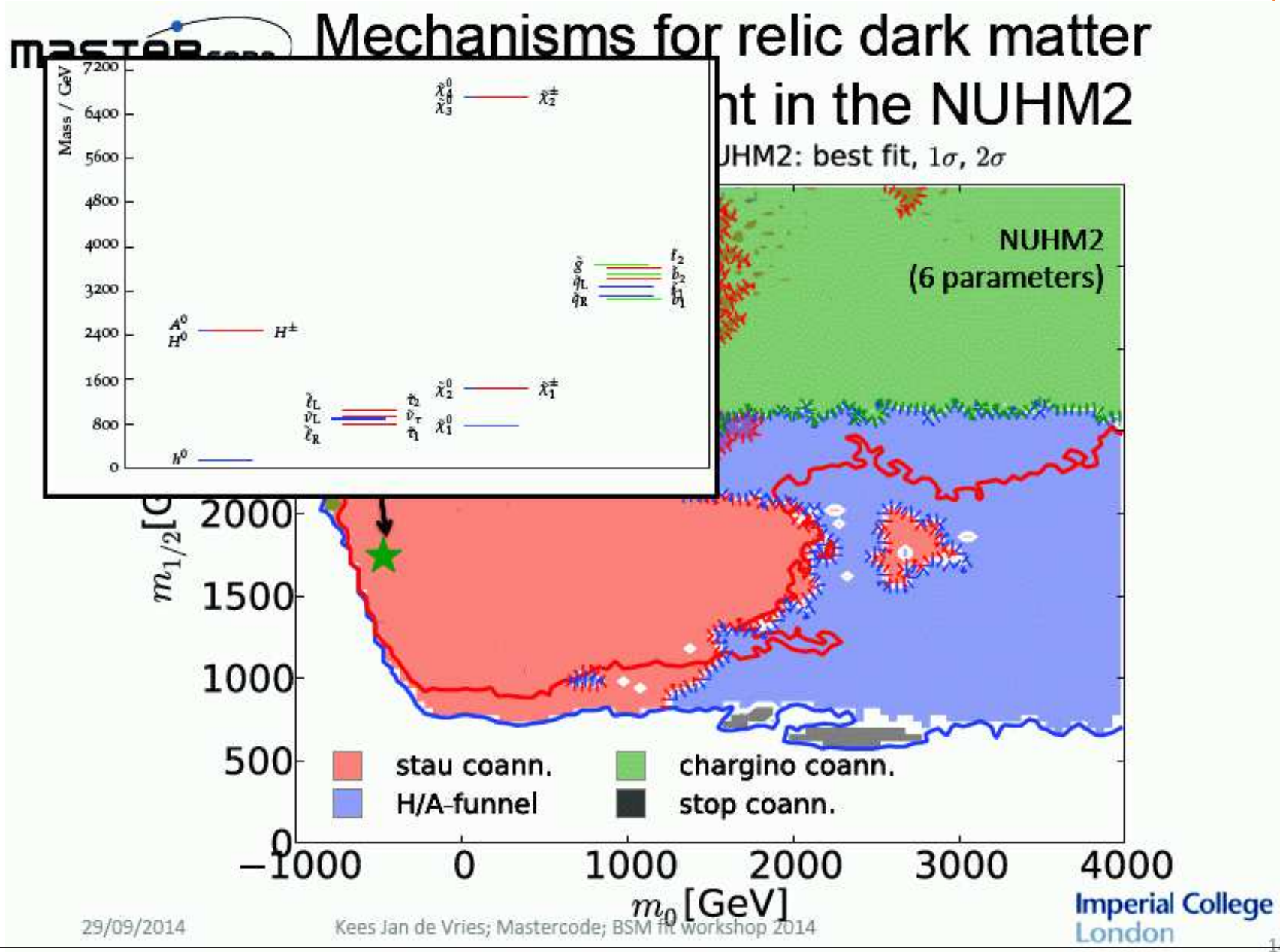


29/09/2014

Kees Jan de Vries; Mastercode; BSM fit workshop 2014

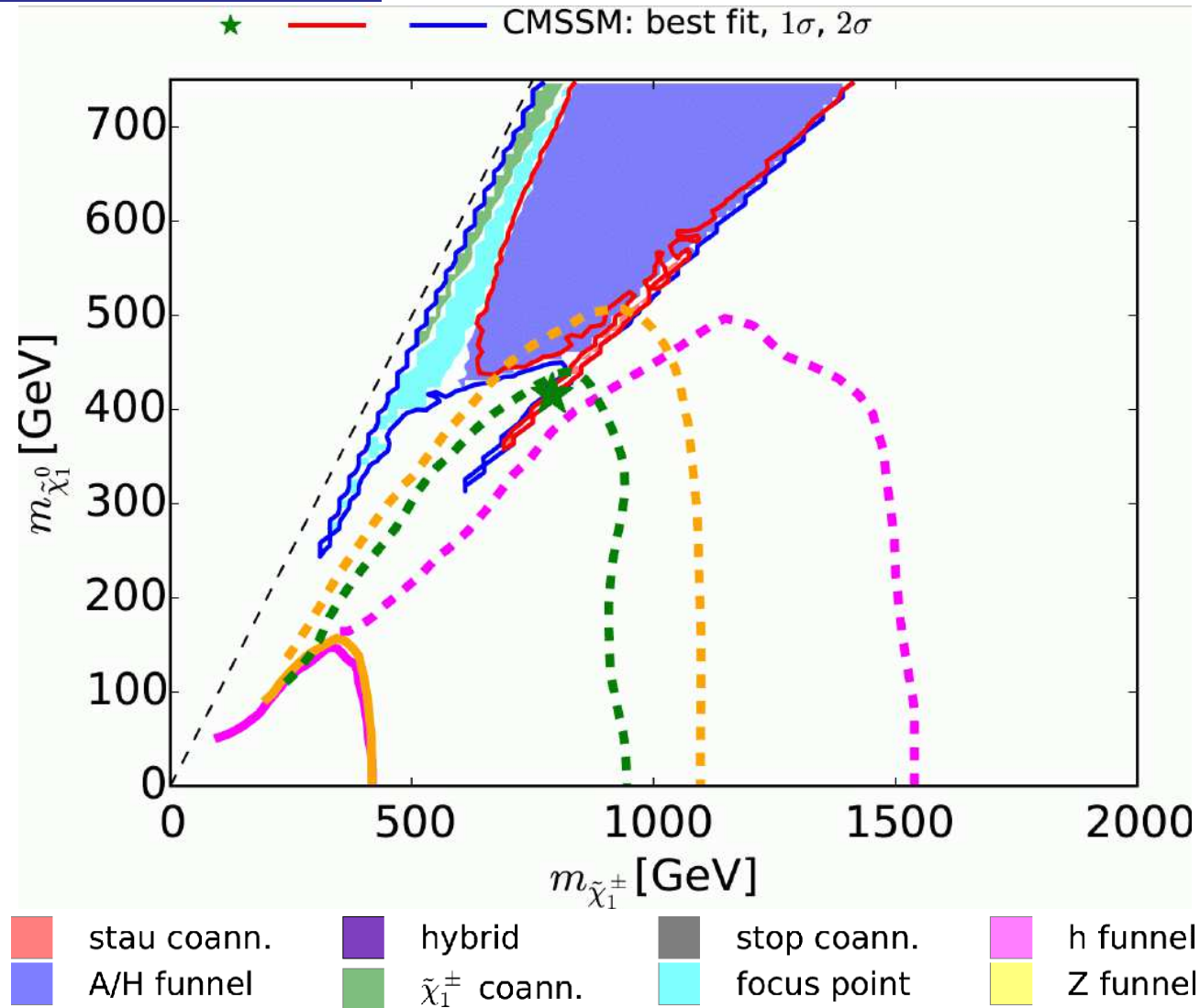
Imperial College
London

13



LHC prospects for CMSSM:

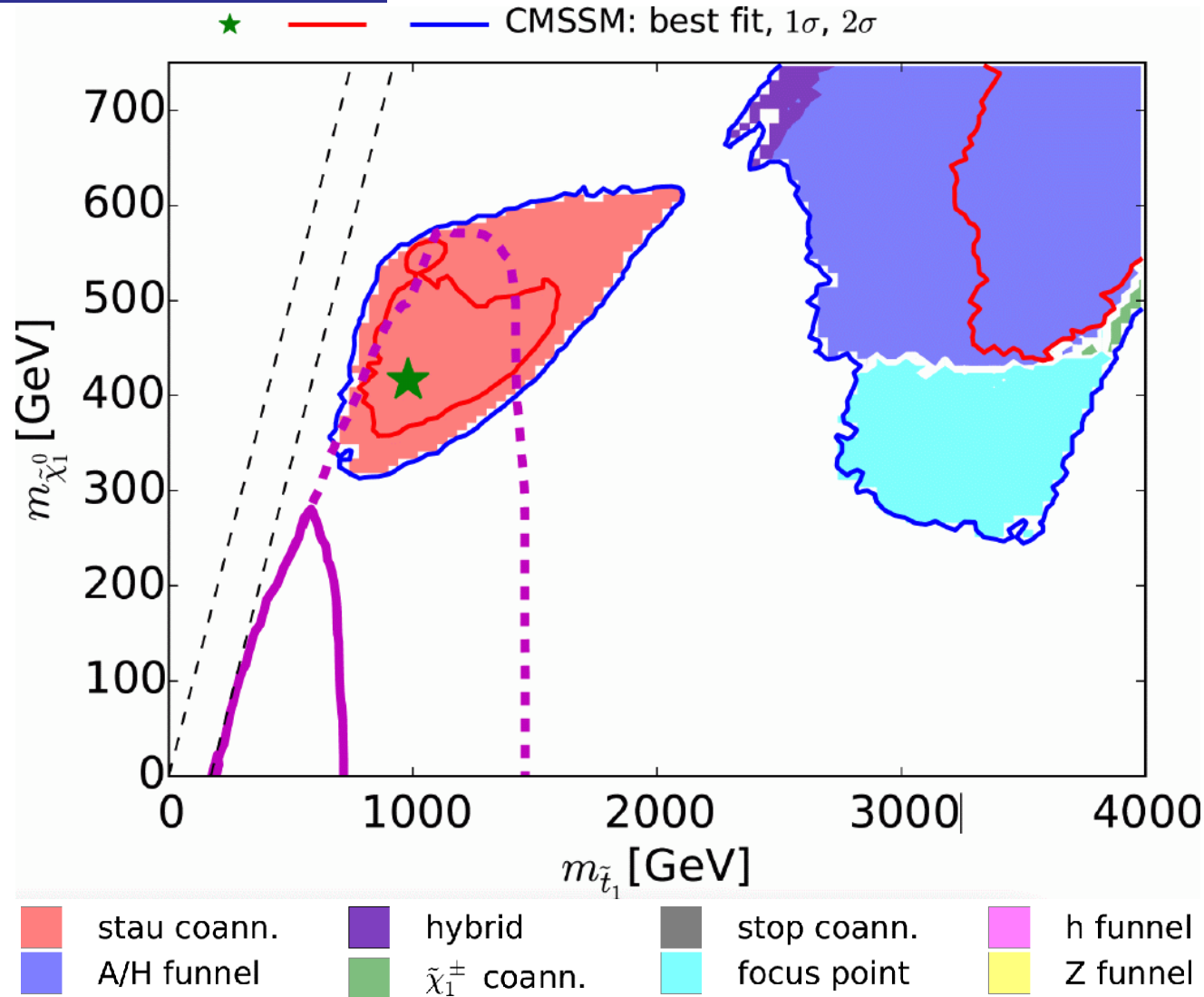
[2015]



solid: current LHC limits, dashed: HL-LHC prospects
 \Rightarrow best-fit regions can be covered! (in EW searches)

LHC prospects for CMSSM:

[2015]



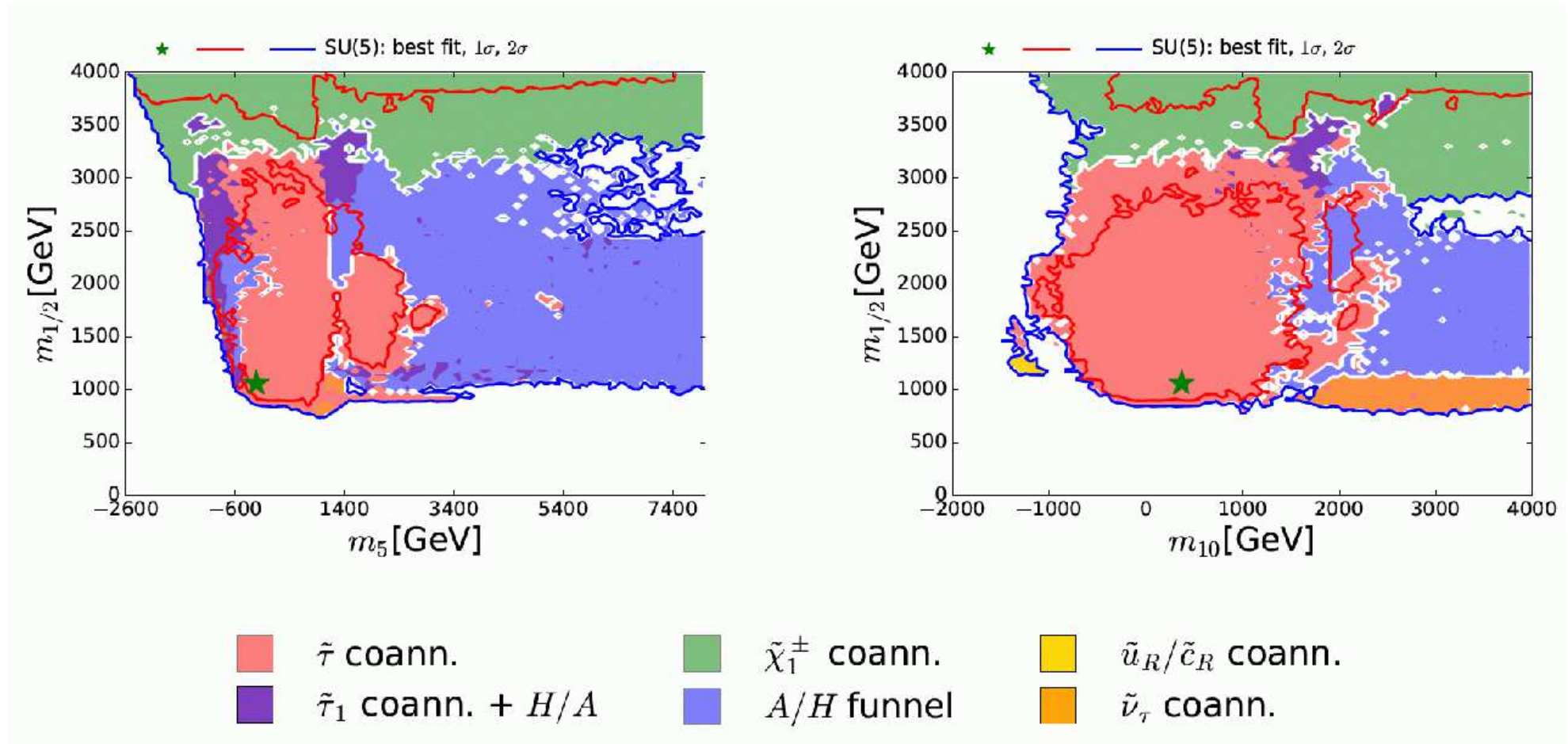
solid: current LHC limits, dashed: HL-LHC prospects
 ⇒ best-fit regions can partially be covered! (in colored searches)

Results in the SU(5)



Dark Matter annihilation mechanism:

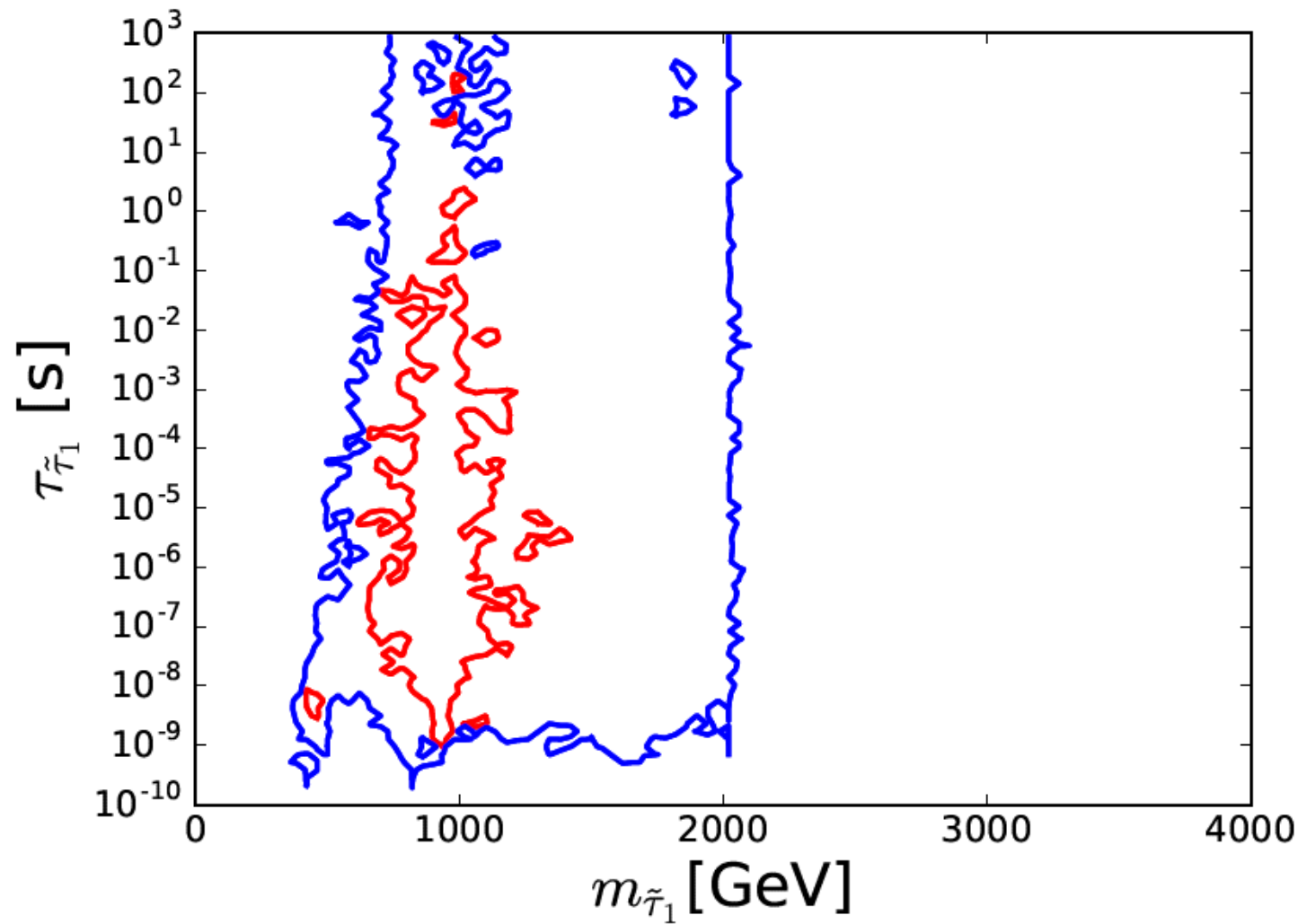
[2016]



$\Rightarrow \tilde{u}_R/\tilde{c}_R/\tilde{\nu}_\tau$ co-ann. possible \Rightarrow but $\tilde{\tau}_1$ co-ann. dominant!

NLSP stau searches at the LHC:

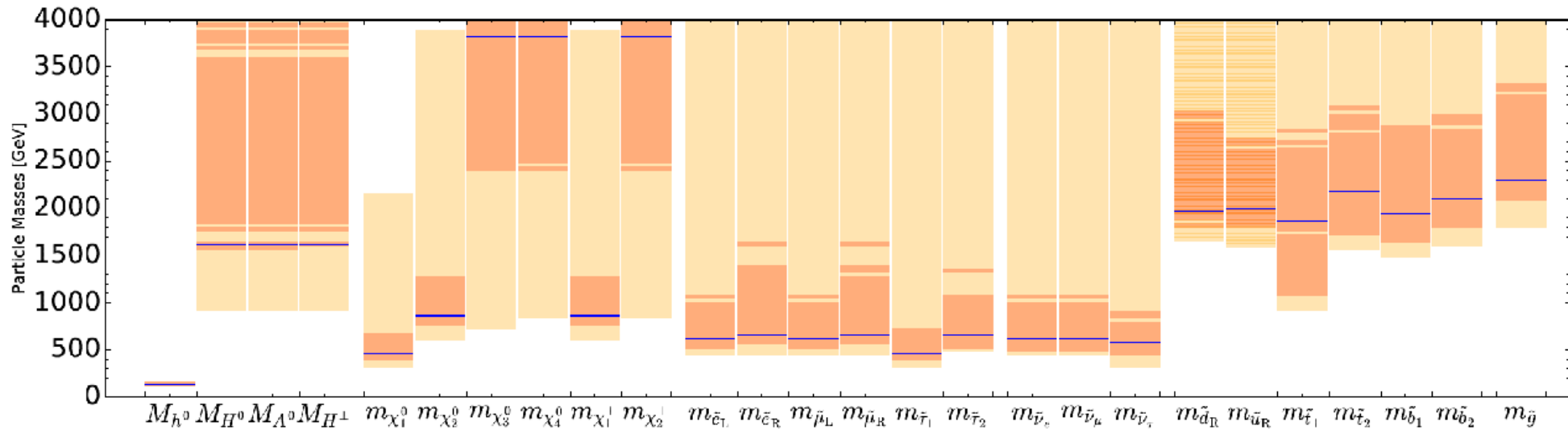
[2016]



⇒ search for long-lived staus!

SU(5) prediction: best-fit masses

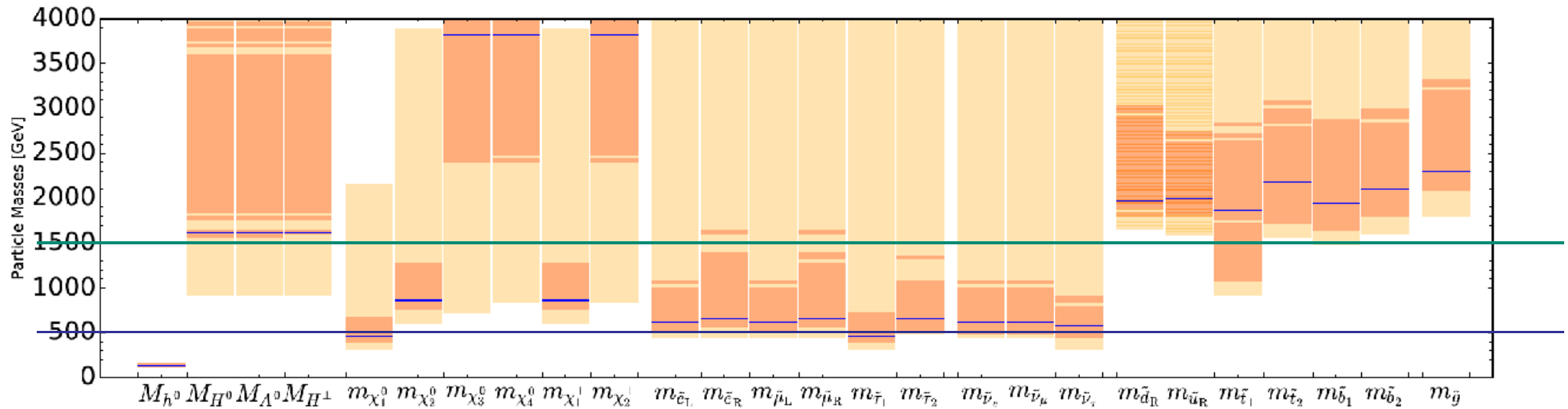
[2016]



- ⇒ high colored masses
- ⇒ lower electroweak masses
 - partially with not too large 1σ ranges
- ⇒ clear prediction for ILC and CLIC

SU(5) prediction: best-fit masses

[2016]



ILC: $\sqrt{s} = 1000$ GeV \Rightarrow only few EW particles possibly accessible

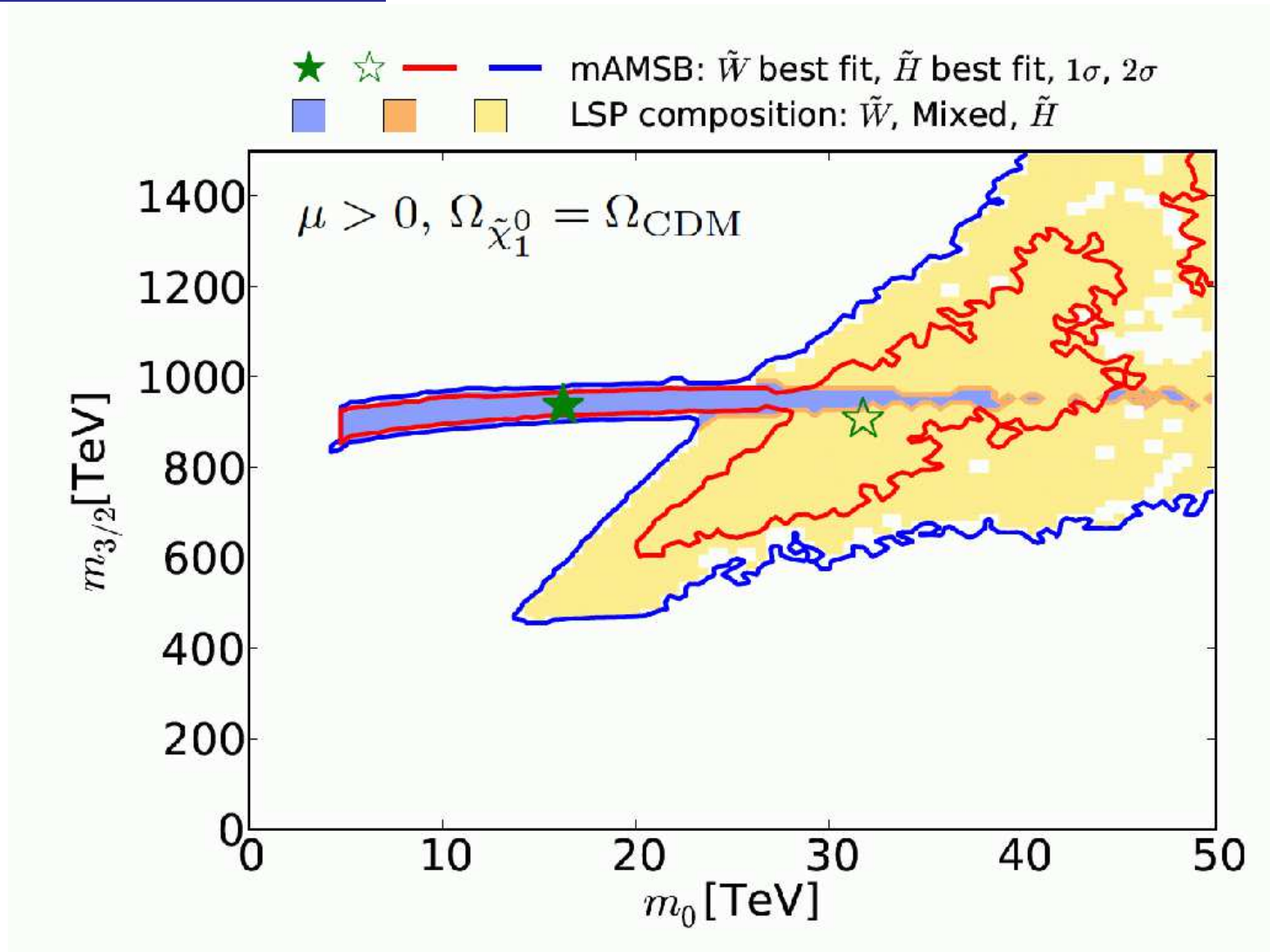
CLIC: $\sqrt{s} = 3000$ GeV \Rightarrow pair production of many SUSY particles “likely”
 \Rightarrow no access to colored particles

Results in the mAMSB



Dark Matter composition:

[2016]



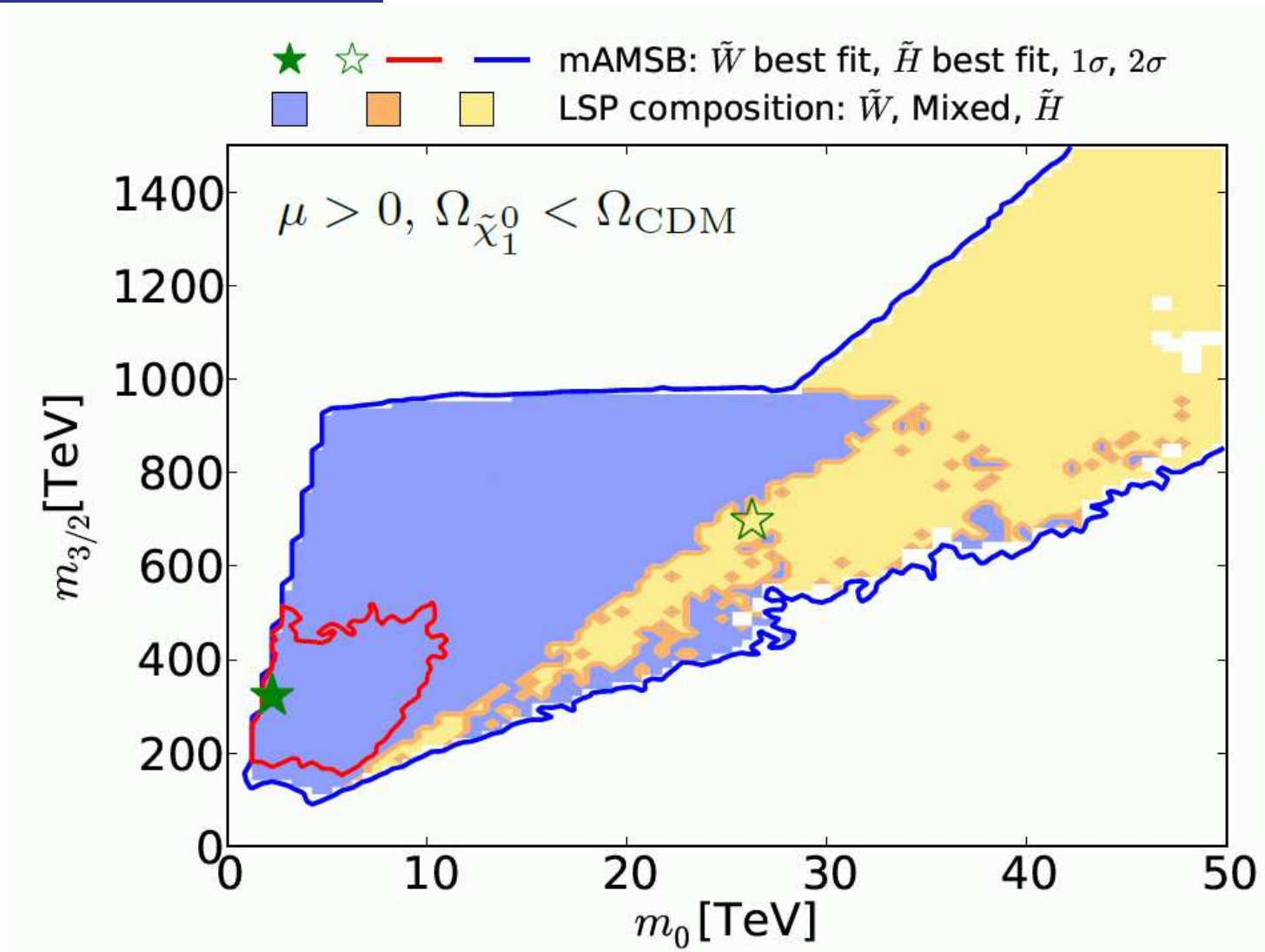
$\Rightarrow m_{\tilde{\chi}_1^0} \sim 2.9 \pm 0.1 \text{ TeV (wino)}, \sim 1.1 \pm 0.02 \text{ TeV (higgsino)}$

Results in the mAMSB



Dark Matter composition:

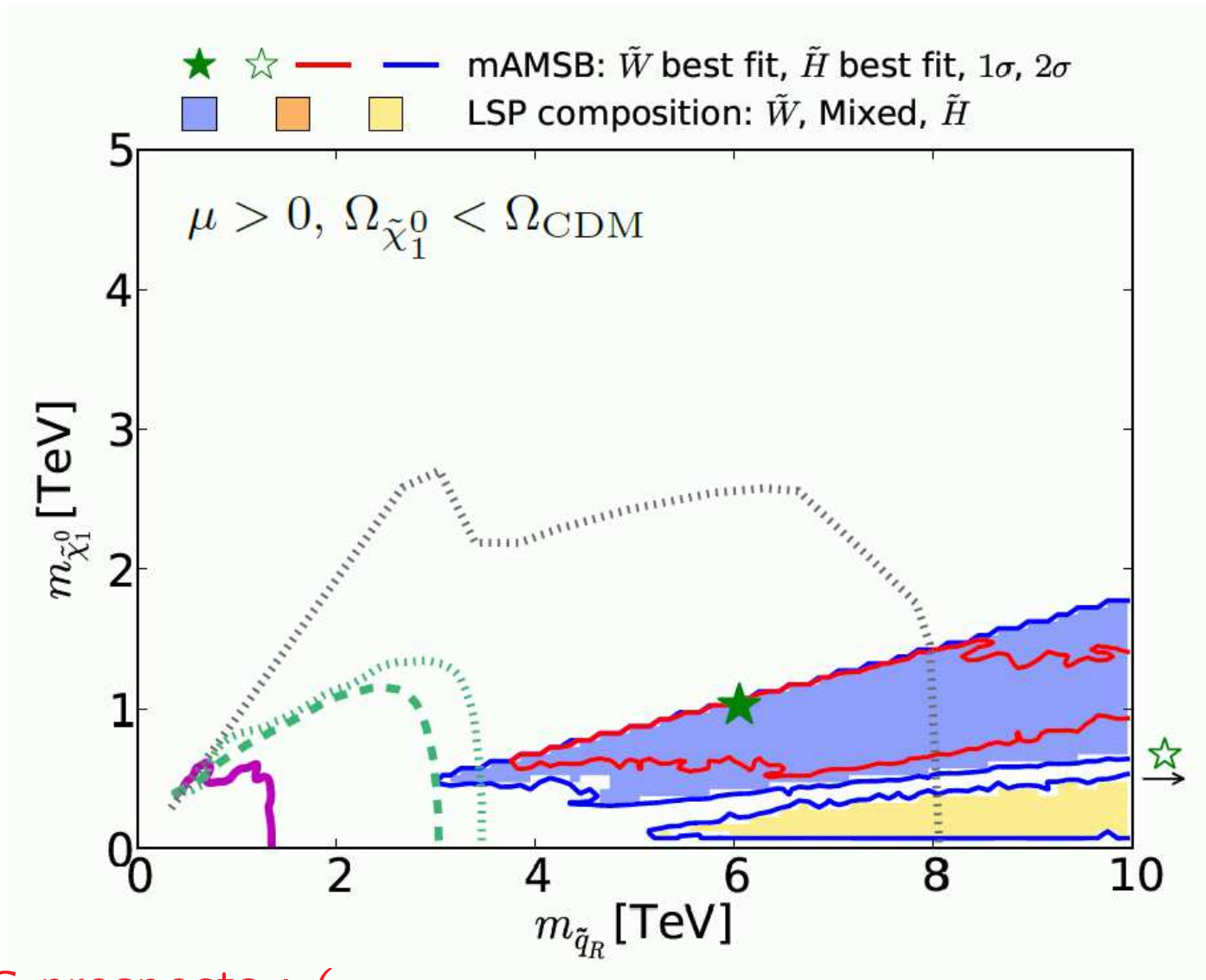
[2016]



⇒ very relaxed limits ⇒ lower masses

Squark mass vs. DM mass:

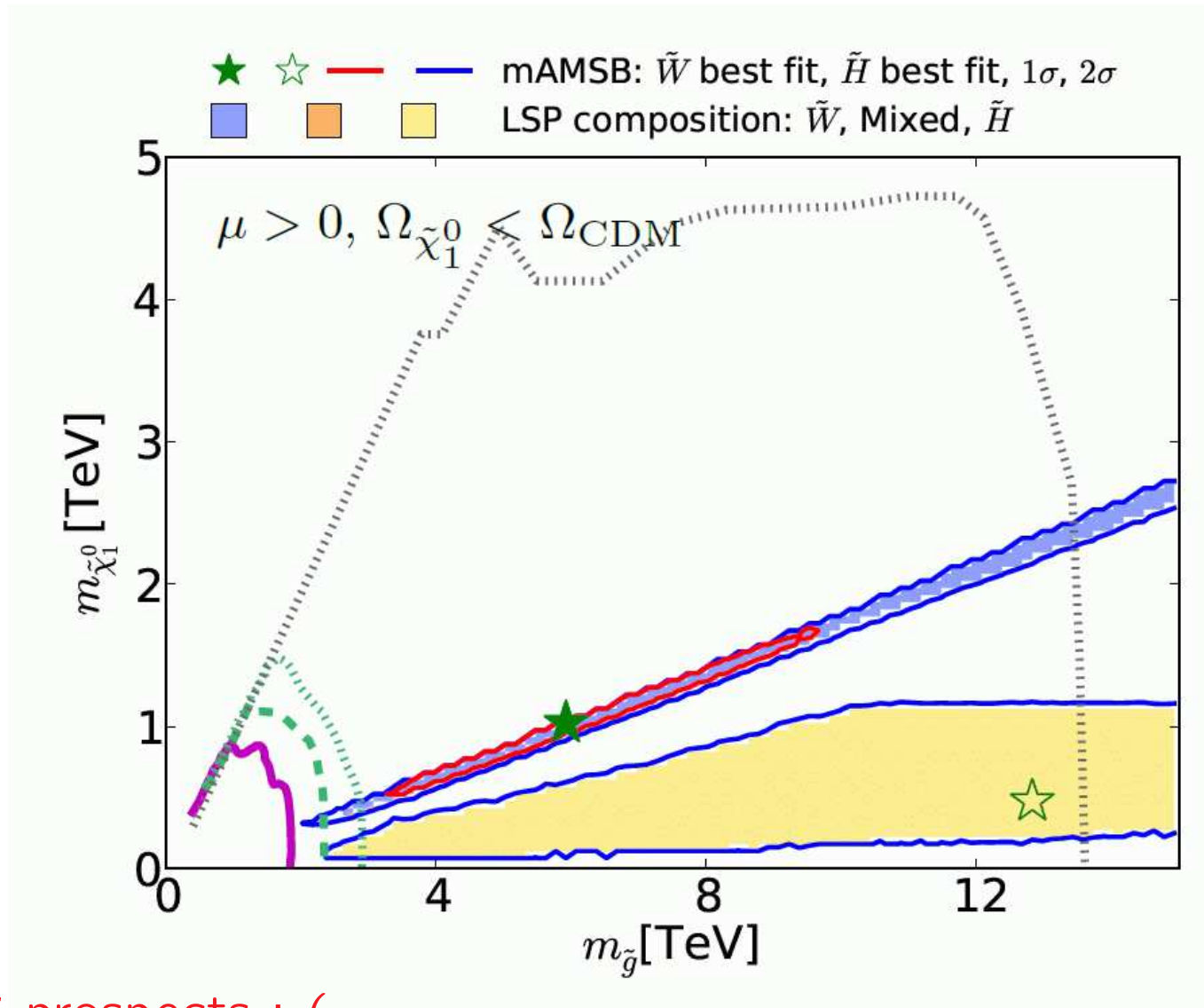
[2016]



⇒ bad LHC prospects :-)

⇒ better FCC-hh prospects :-)

Gluino mass vs. DM mass:

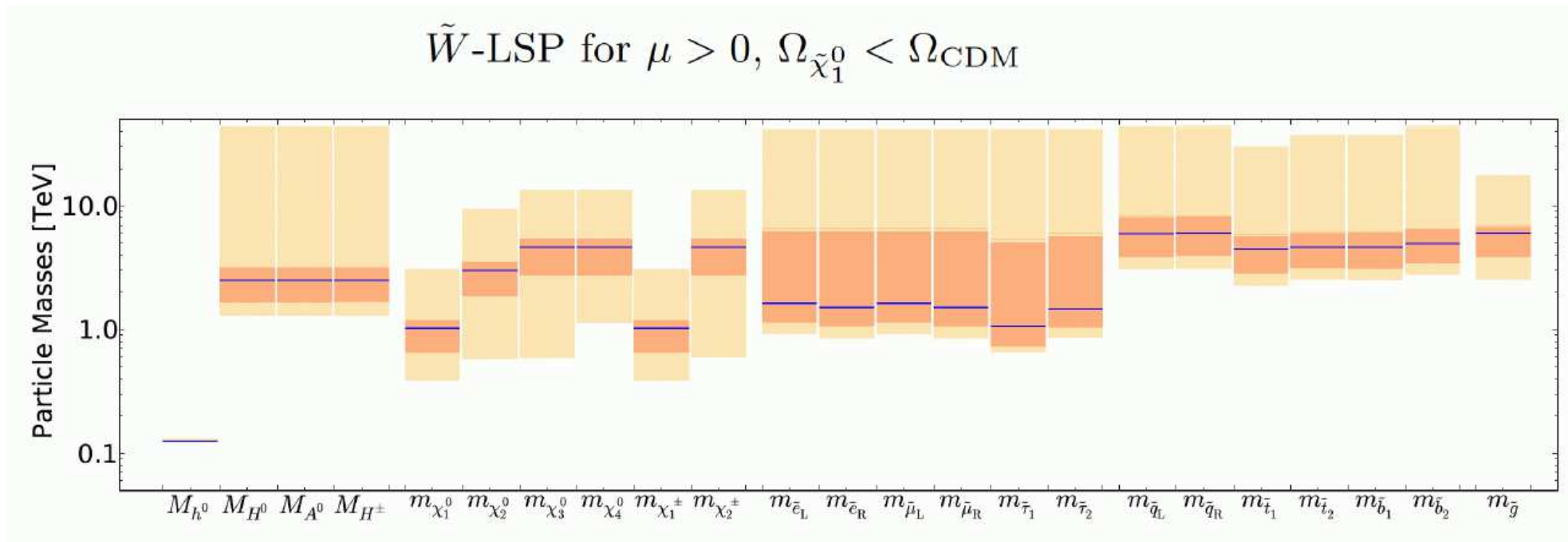


⇒ bad LHC prospects :-)

⇒ better FCC-hh prospects :-)

mAMSB prediction: best-fit masses (wino)

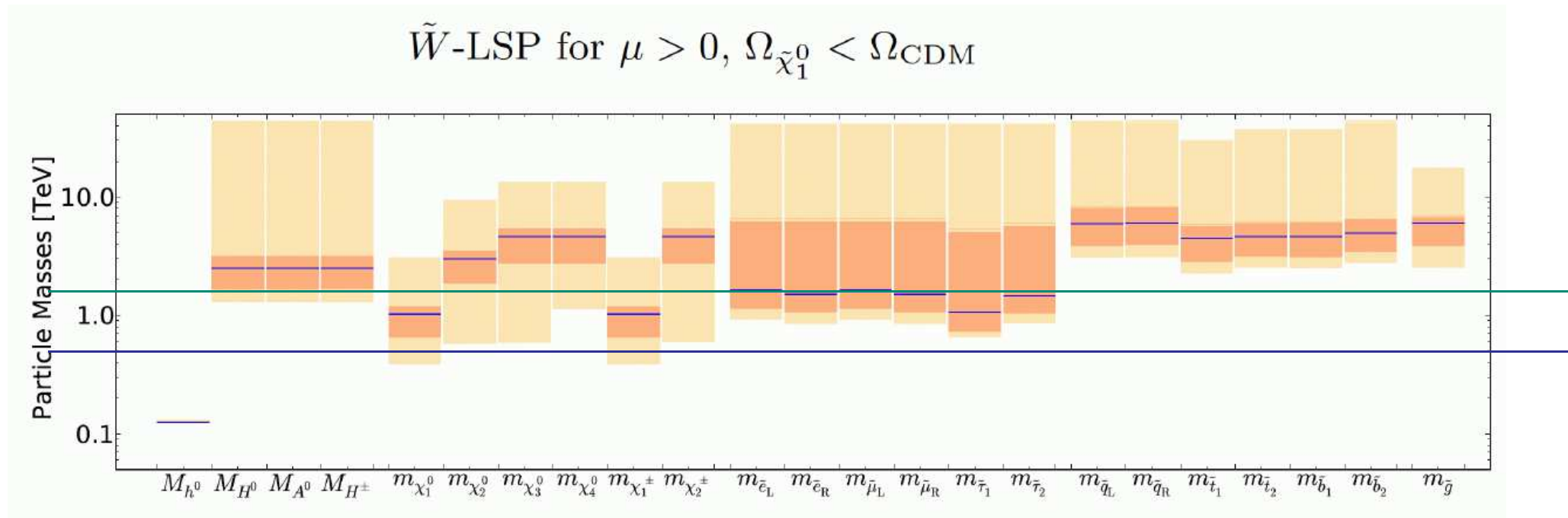
[2016]



- ⇒ high colored masses
- ⇒ lower electroweak masses
 - partially with not too large 1σ ranges
- ⇒ clear prediction for ILC and CLIC

mAMSB prediction: best-fit masses (wino)

[2016]

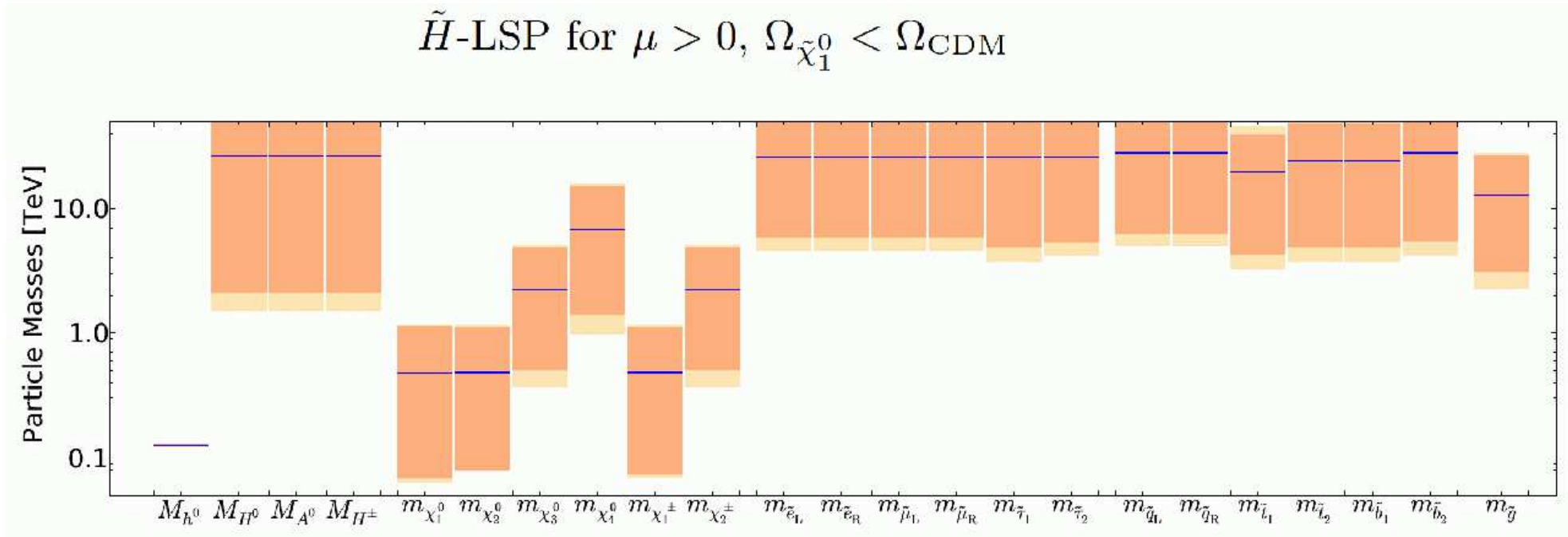


ILC: $\sqrt{s} = 1000 \text{ GeV} \Rightarrow$ bad prospects

CLIC: $\sqrt{s} = 3000 \text{ GeV} \Rightarrow$ pair production of few SUSY particles “likely”
 \Rightarrow no access to colored particles

mAMSB prediction: best-fit masses (higgsino)

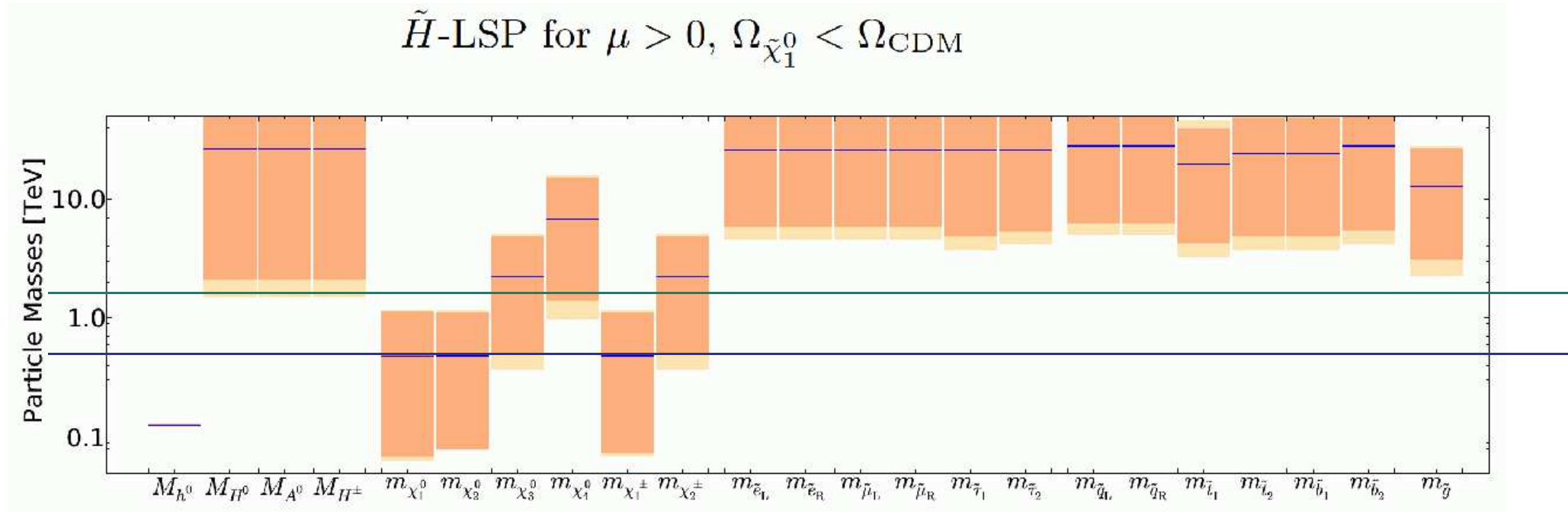
[2016]



- ⇒ high colored masses
- ⇒ some(!) lower electroweak masses
 - partially with not too large 2σ ranges
- ⇒ clear prediction for ILC and CLIC

mAMSB prediction: best-fit masses (higgsino)

[2016]



ILC: $\sqrt{s} = 1000$ GeV \Rightarrow few EW particles possibly accessible

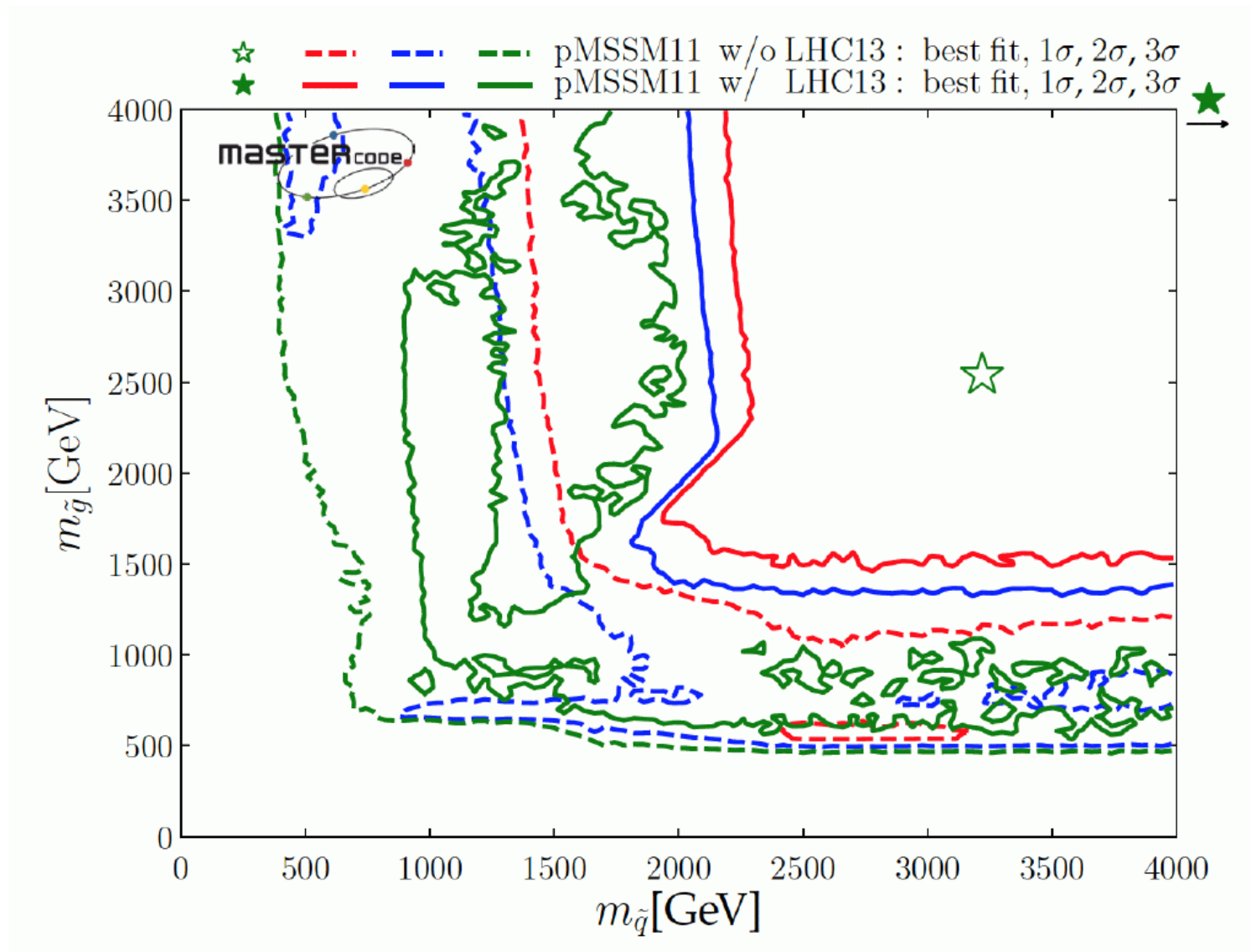
CLIC: $\sqrt{s} = 3000$ GeV \Rightarrow pair production of few SUSY particles
 “guraranteed”
 \Rightarrow no access to colored particles

3. Results and predictions in the pMSSM11

Parameter	Range	Number of segments
M_1	(-4 , 4) TeV	6
M_2	(0 , 4) TeV	2
M_3	(-4 , 4) TeV	4
$m_{\tilde{q}}$	(0 , 4) TeV	2
$m_{\tilde{q}_3}$	(0 , 4) TeV	2
$m_{\tilde{l}}$	(0 , 2) TeV	1
$m_{\tilde{\tau}}$	(0 , 2) TeV	1
M_A	(0 , 4) TeV	2
A	(-5 , 5) TeV	1
μ	(-5 , 5) TeV	1
$\tan \beta$	(1 , 60)	1
Total number of boxes		384

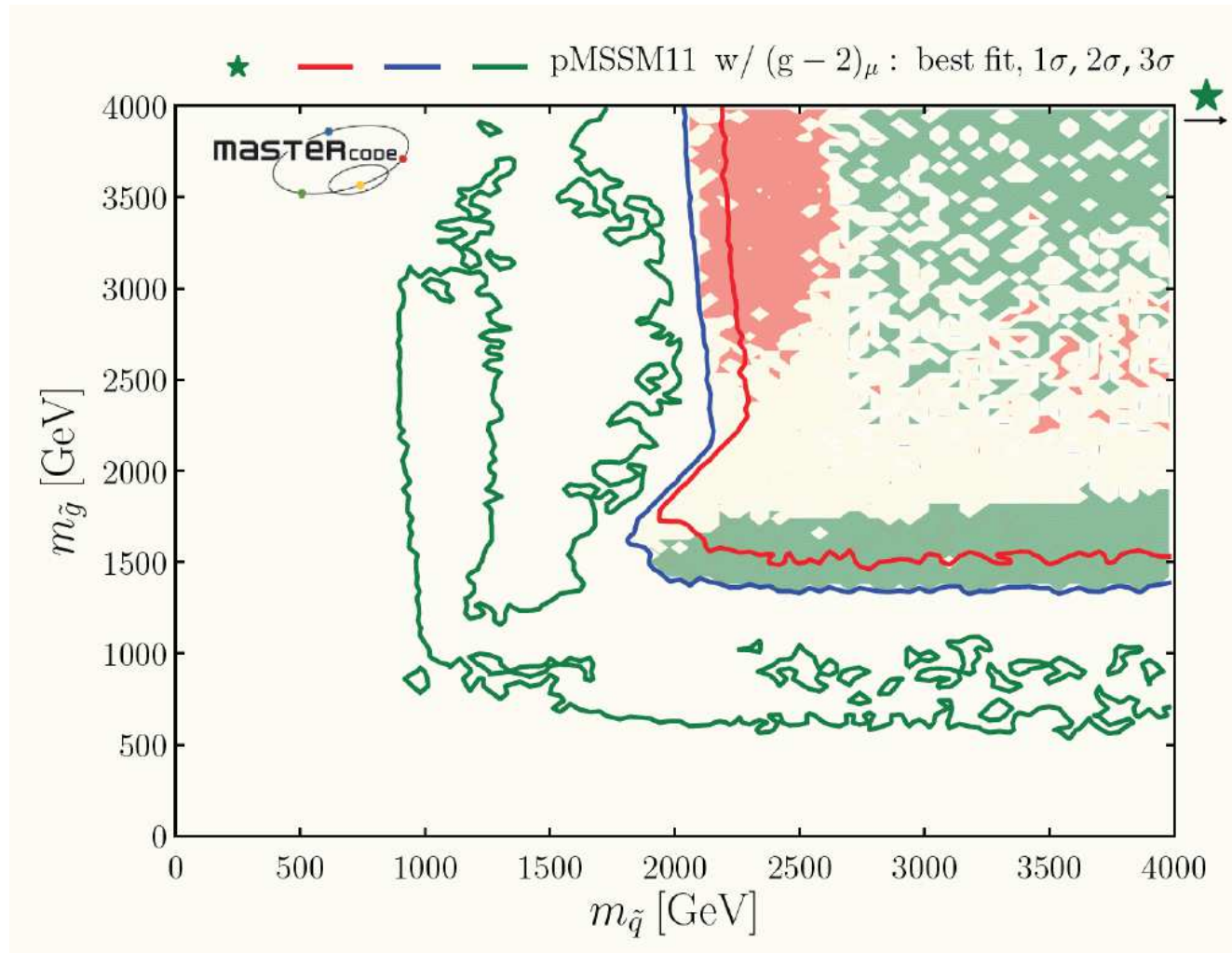
pMSSM11: Going from 8 TeV to 13 TeV (and adding latest DM limits)

[2017]



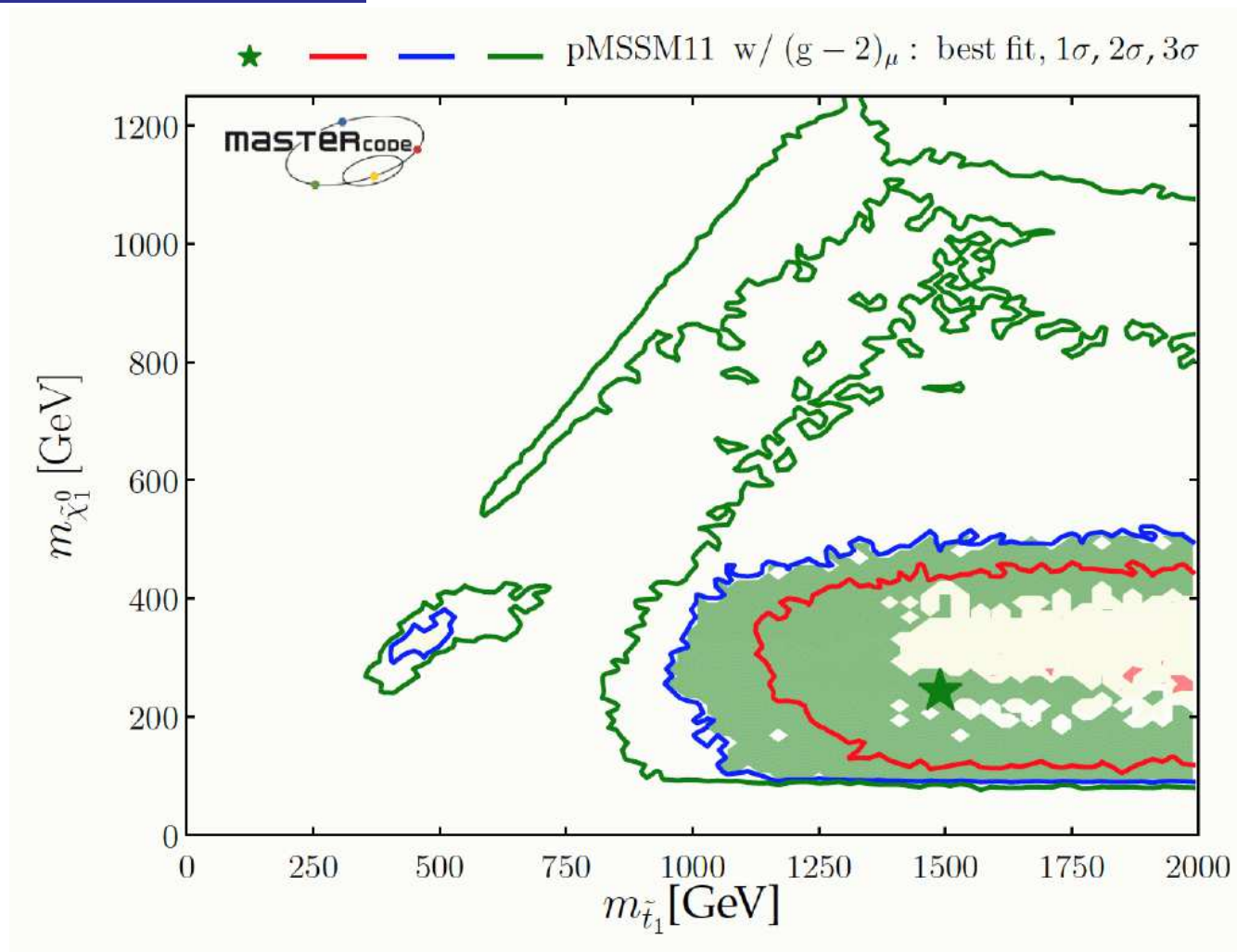
⇒ substantial move to higher masses!

⇒ notice the “nose”!

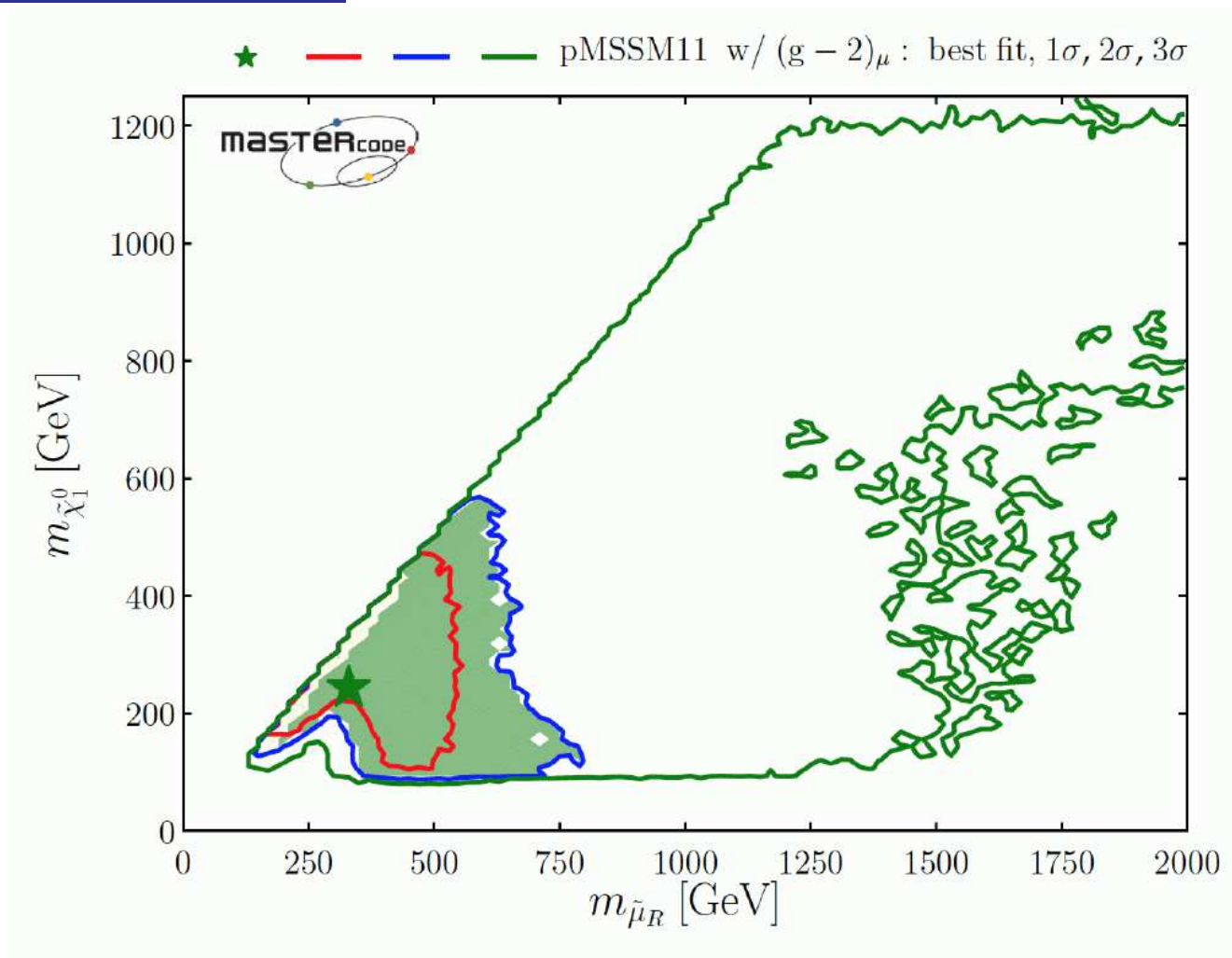


⇒ high masses favored

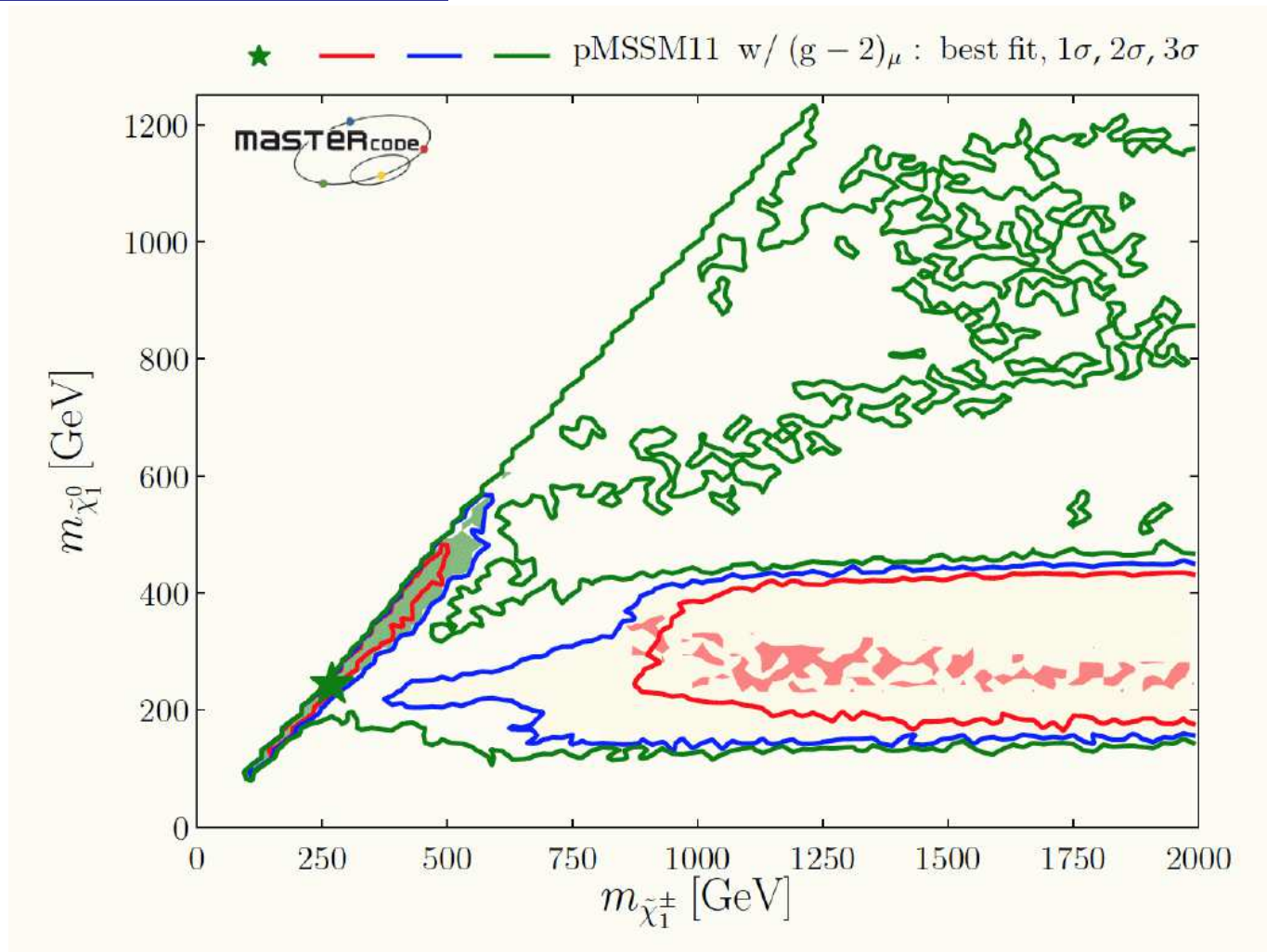
⇒ notice the “nose”!



\Rightarrow high (low) stop (neutralino) masses \Rightarrow notice the compressed region!

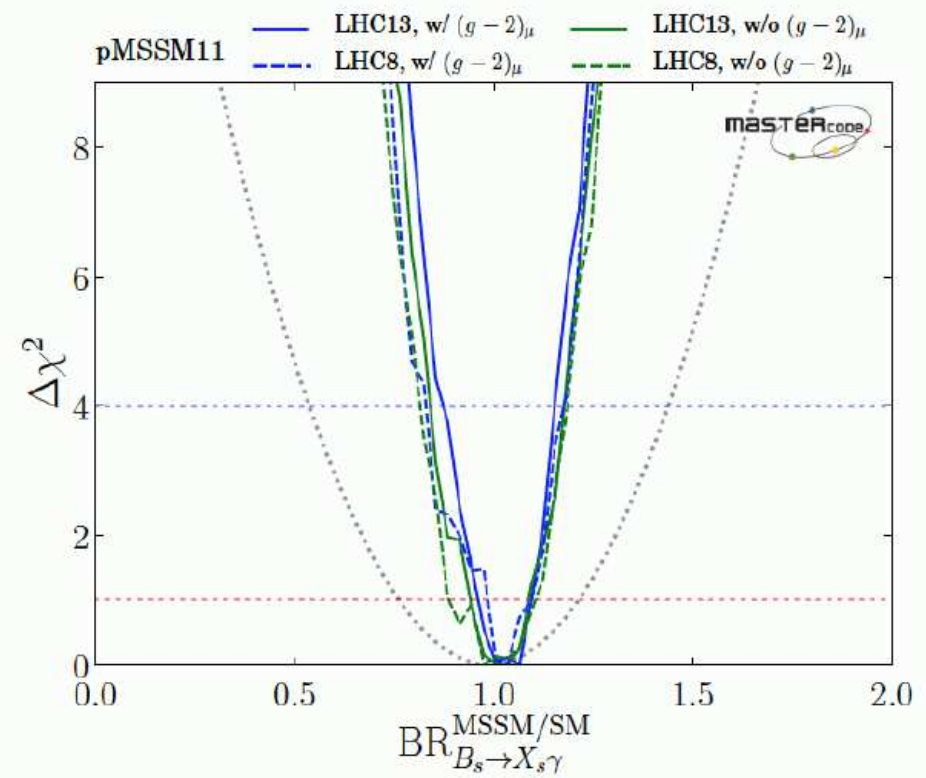
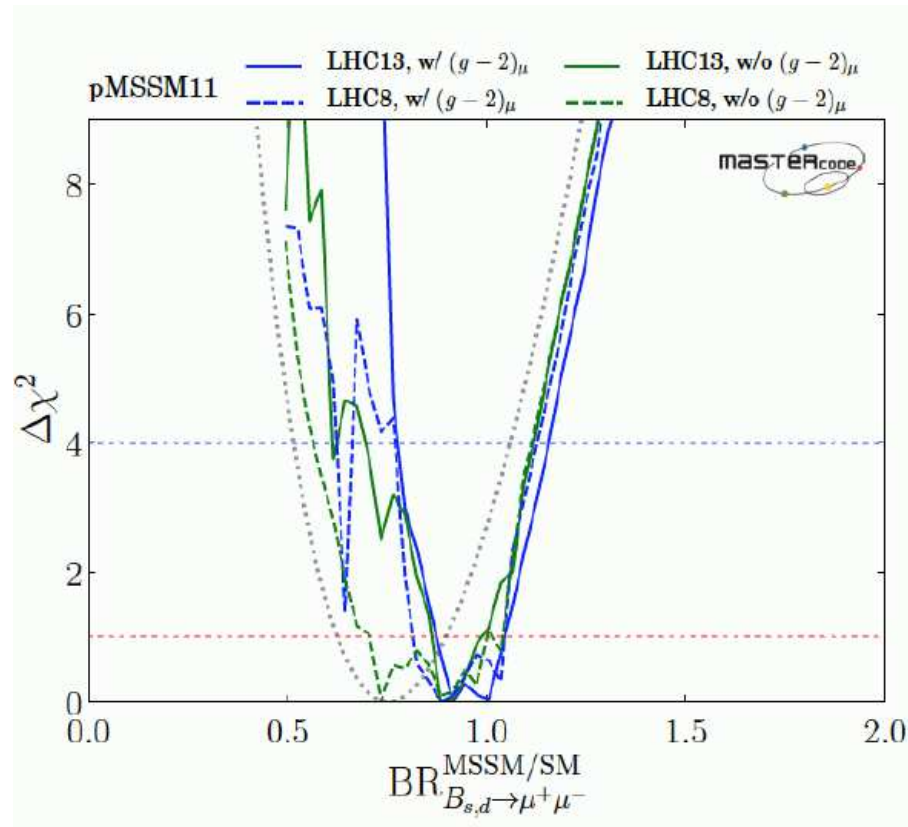


⇒ all masses low!!



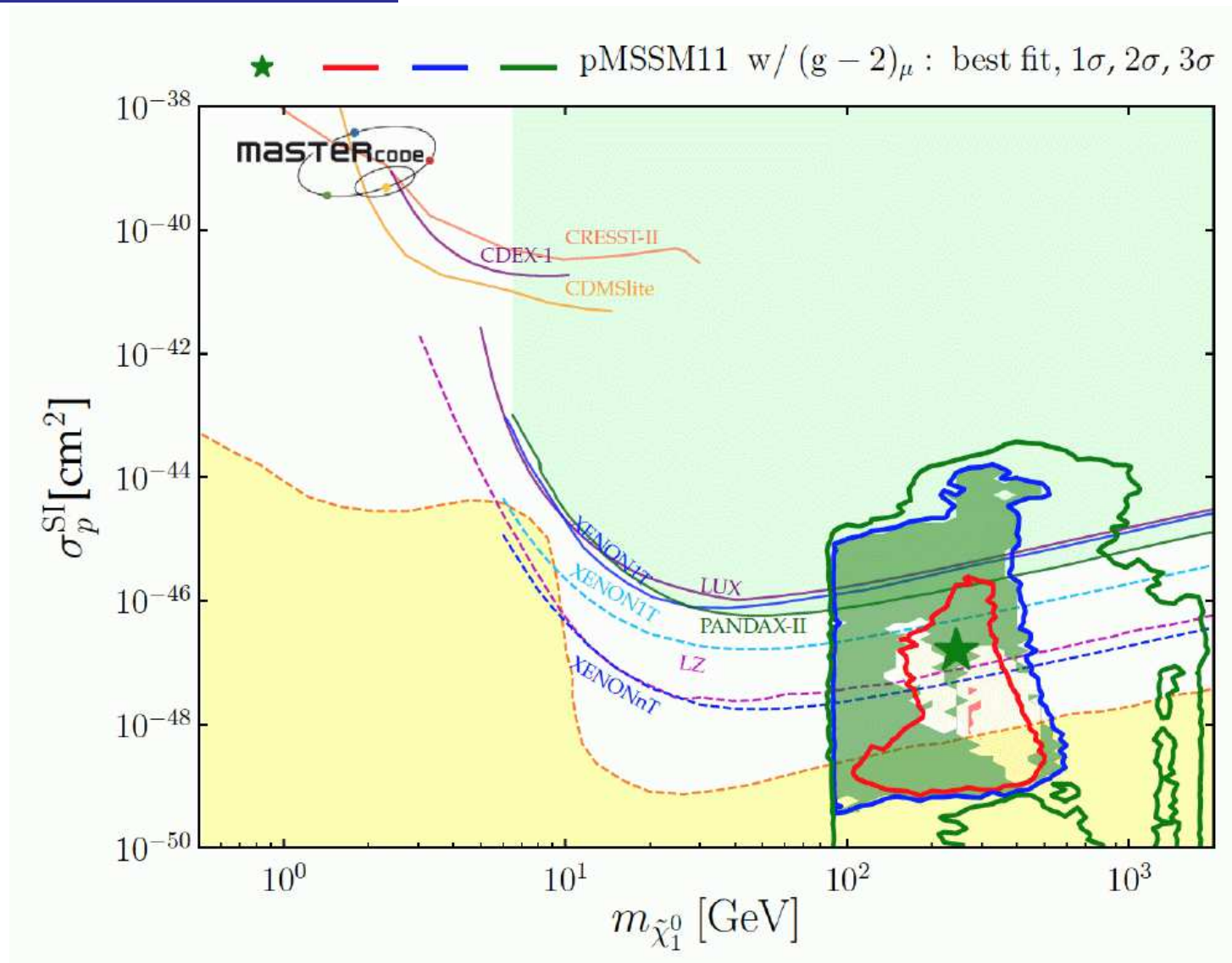
⇒ chargino co-annihilation

⇒ $M_1 \sim M_2$



⇒ follows the experimental data

⇒ $BR(B_s \rightarrow \mu^+ \mu^-)$: below the SM value



⇒ good prospects for future experiments

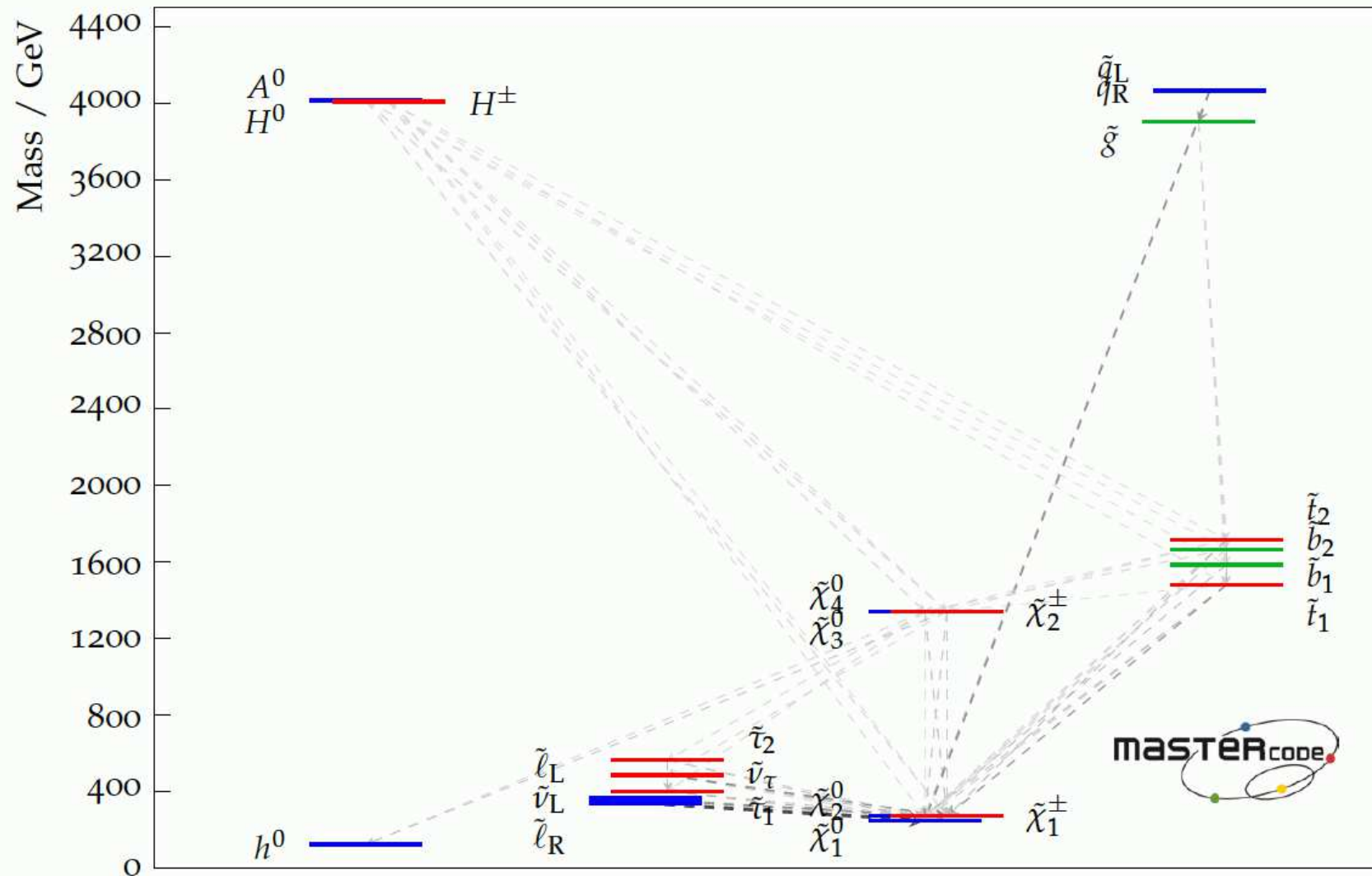
⇒ but no guarantee

pMSSM11: best-fit point parameters

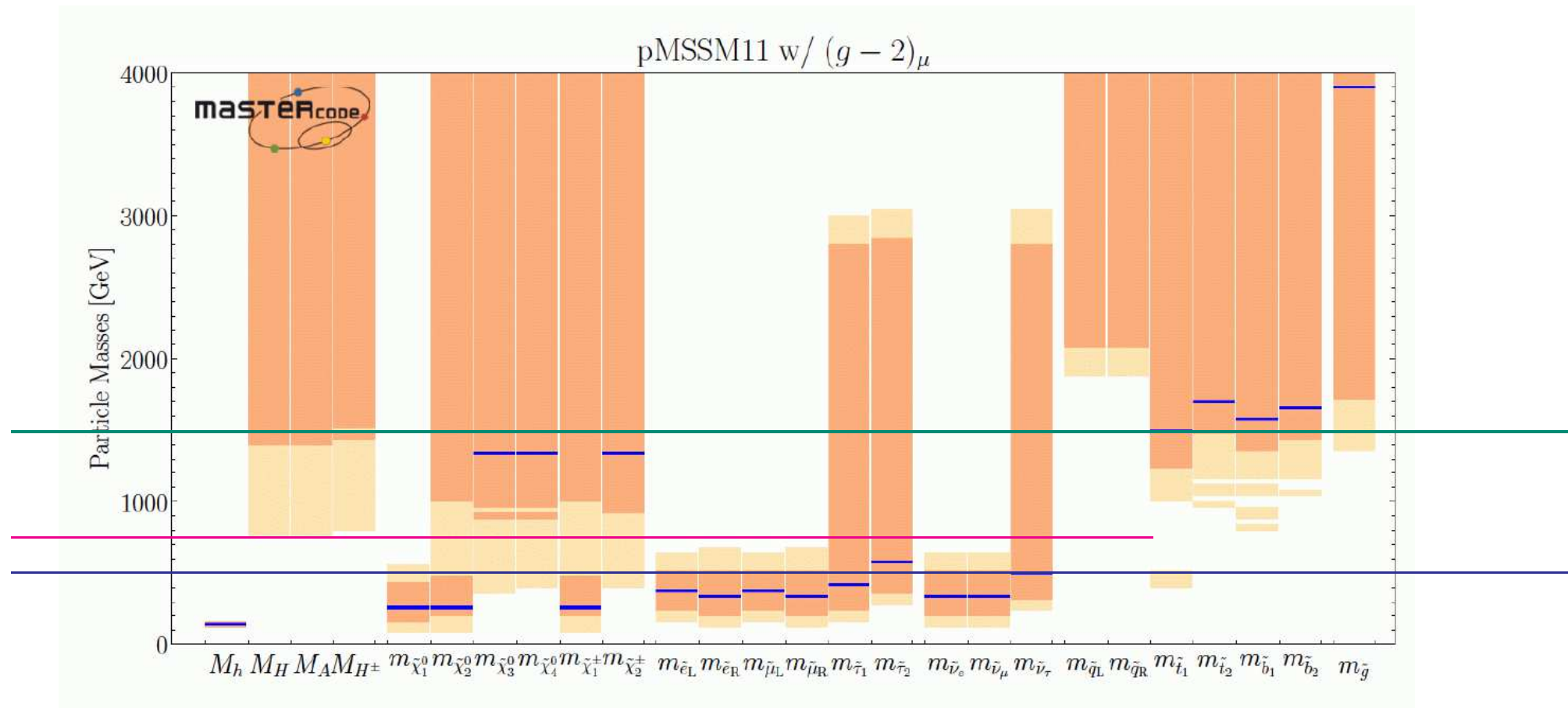
[2017]

Parameter	With LHC 13 TeV and $(g - 2)_\mu$	
	Best fit	'Nose' region
M_1	0.25 TeV	- 0.39 TeV
M_2	0.25 TeV	1.2 TeV
M_3	- 3.86 TeV	- 1.7 TeV
$m_{\tilde{q}}$	4.0 TeV	2.00 TeV
$m_{\tilde{q}_3}$	1.7 TeV	4.1 TeV
$m_{\tilde{\ell}}$	0.35 TeV	0.36 TeV
$m_{\tilde{\tau}}$	0.46 TeV	1.4 TeV
M_A	4.0 TeV	4.2 TeV
A	2.8 TeV	5.4 TeV
μ	1.33 TeV	- 5.7 TeV
$\tan \beta$	36	19
$\chi^2/\text{d.o.f.}$	22.1/20	24.46/20
p-value	0.33	0.22
$\chi^2(HS)$	68.01	67.97

⇒ excellent p value!



⇒ heavy colored, light uncolored spectrum



ILC: $\sqrt{s} = 1000$ GeV \Rightarrow precision analysis of EW particle and DM easy!

ILC: $\sqrt{s} = 1000$ GeV \Rightarrow higher reach for non-diagonal production!

CLIC: $\sqrt{s} = 3000$ GeV \Rightarrow precision analysis of EW particles and DM easy!

What to conclude?

What to conclude?

⇒ **Look at the p values!**

What to conclude?

⇒ **Look at the p values!**

Model	Min. χ^2/dof	χ^2 -prob. (p -value)
CMSSM	32.8/18	11%
NUHM1	31.1/23	12%
NUHM2	30.3/22	11%
SU(5)	32.4/23	9%
mAMSB	36.5/27	11%
pMSSM11	21.0/20	33%

Which model is more likely??

What to conclude?

⇒ **Look at the p values!**

Model	Min. χ^2/dof	χ^2 -prob. (p -value)
CMSSM	32.8/18	11%
NUHM1	31.1/23	12%
NUHM2	30.3/22	11%
SU(5)	32.4/23	9%
mAMSB	36.5/27	11%
pMSSM11	21.0/20	33%

Which model is more likely??

⇒ pMSSM11: model with higher χ^2 -probability
model with interesting ILC prospects
model with good CLIC prospects
detailed LHC analysis tbd!

4. $\phi_{96} \rightarrow \gamma\gamma$

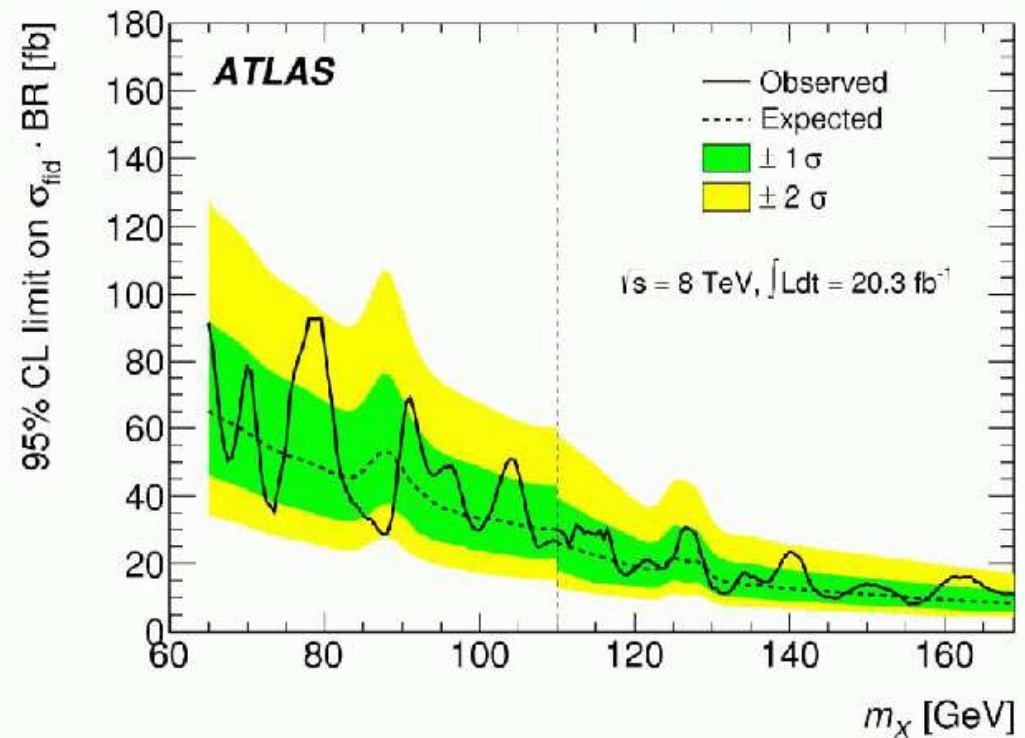
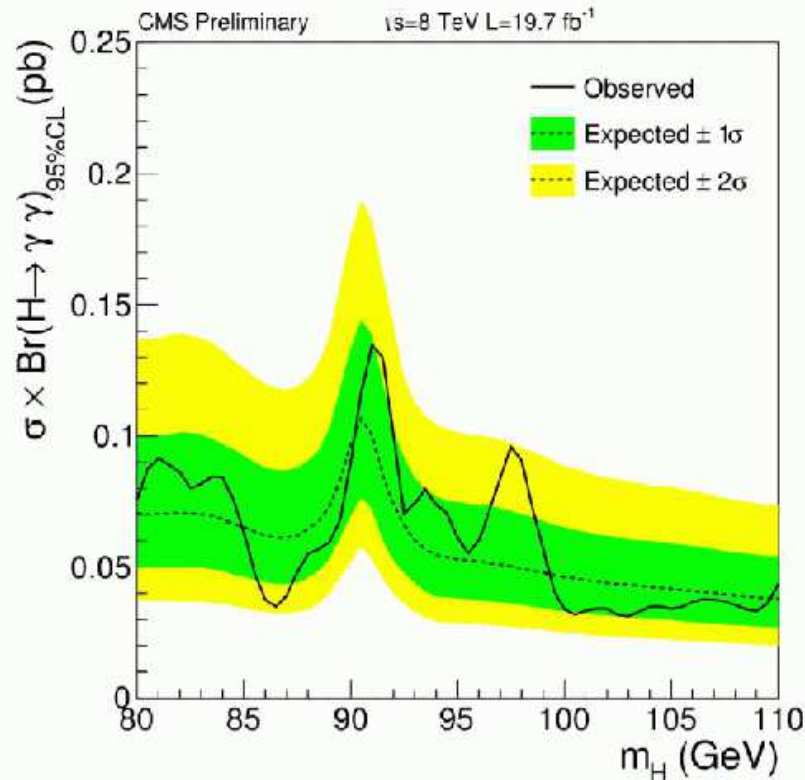
- What was seen in Run I?
- What was seen in Run II?
- What was seen at LEP?
- Should we get excited?
- Which model fits?



CMS PAS HIG-14-037

$h \rightarrow \gamma\gamma$ (65-110 GeV) Run 1

PRL 113 171801 (2014)



• $\sim 2\sigma$ excursion @ ~ 97.5 GeV

• $\sim 2\sigma$ excursion @ ~ 80 GeV

18

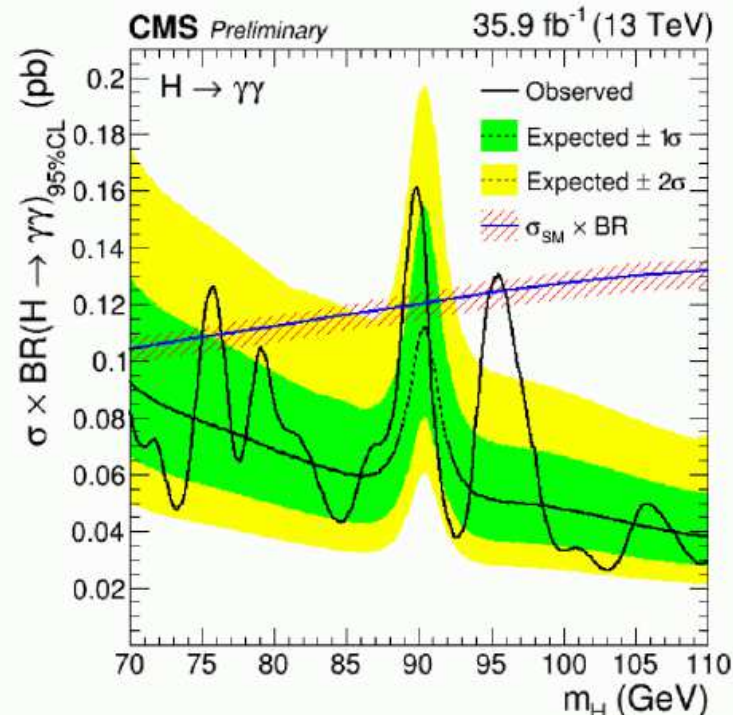
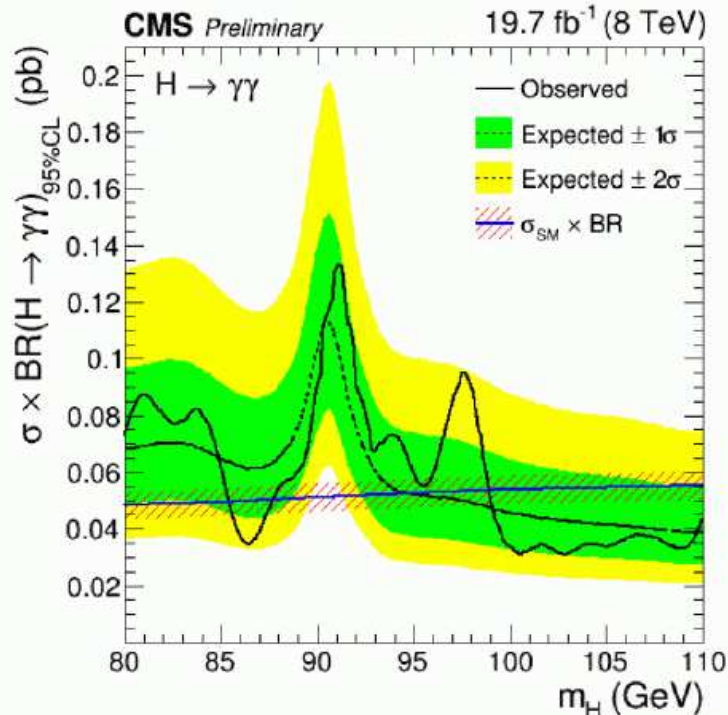
S. Gascon-Shotkin HDays17, Santander, ES Sept. 22 2017



$h \rightarrow \gamma\gamma$ (70-110 GeV) Runs 1+2



CMS PAS HIG-17-013



8 TeV:
 minimum(maximum)
 limit on $\sigma \times Br$:
 31(133) fb at
 $m=102.8(91.1)$ GeV

13 TeV:
 minimum(maximum)
 limit on $\sigma \times Br$:
 26(161) fb at
 $m=103.0(89.9)$ GeV

- 8 TeV limits on $\sigma \times Br$ redone with 0.1 GeV step. Production processes assumed in SM proportions. No significant excess with respect to expected limits observed.

26

S. Gascon-Shotkin HDays17, Santander, ES Sept. 22 2017

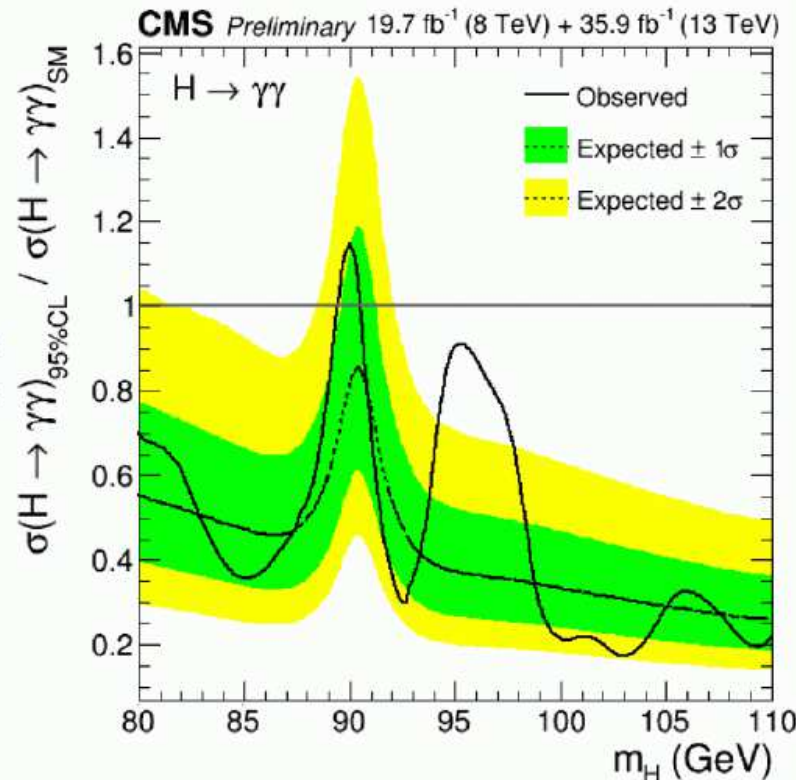


$h \rightarrow \gamma\gamma$ (70-110 GeV) Runs 1+ 2



CMS PAS HIG-17-013

All experimental + theoretical systematic uncertainties assumed uncorrelated except for those on signal acceptance due to scale variations + those on production cross sections (assumed 100% correlated).



8 TeV+13 TeV:
 minimum(maximum) limit
 on $(\sigma \times \text{Br}) / (\sigma \times \text{Br})_{\text{SM}}$:
 0.17(1.15) at
 $m=103.0(90.0)\text{GeV}$

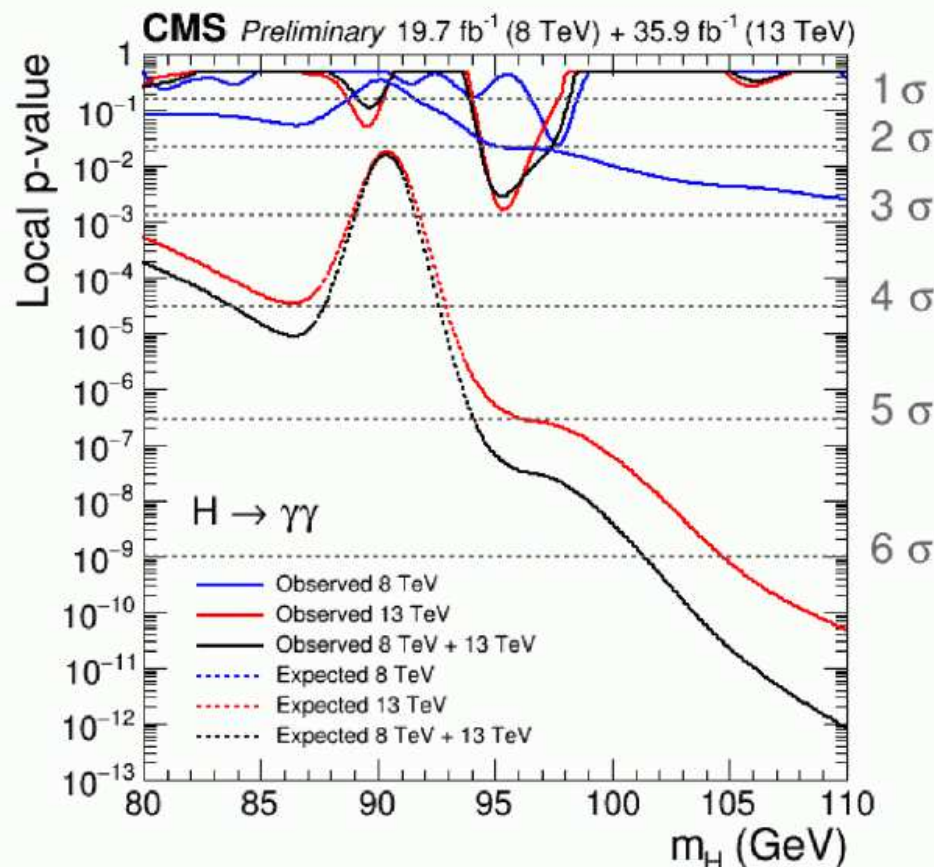
- Combined 8 TeV+13 TeV $\sigma \times \text{BR}$ limit normalized to SM expectation (production processes assumed in SM proportions). No significant excess with respect to expected limits observed.

29

S. Gascon-Shotkin HDays17, Santander, ES Sept. 22 2017



$h \rightarrow \gamma\gamma$ (70-110 GeV) Runs 1+ 2



CMS PAS HIG-17-013

8 TeV: Excess with $\sim 2.0 \sigma$ local significance at $m=97.6$ GeV

13 TeV: Excess with $\sim 2.9 \sigma$ local (1.47σ global) significance at $m=95.3$ GeV

8TeV+13 TeV: Excess with $\sim 2.8 \sigma$ local (1.3σ global) significance at $m=95.3$ GeV

More data are required to ascertain the origin of this excess

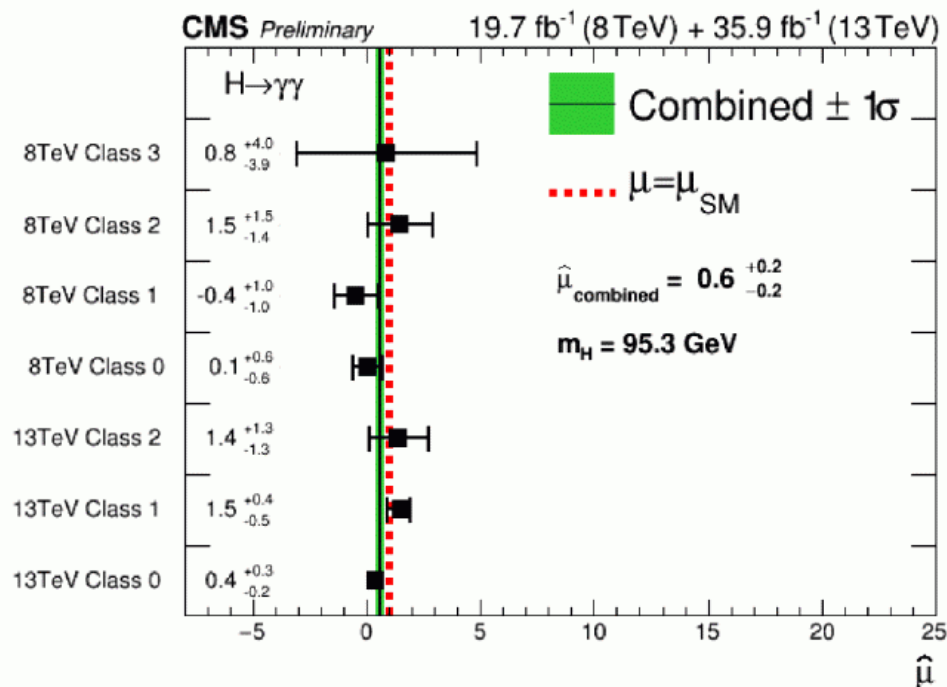
- Expected and observed local p-values for 8 TeV, 13 TeV and their combination

S. Gascon-Shotkin HDays17, Santander, ES Sept. 22 2017

30



$h \rightarrow \gamma\gamma$ (70-110 GeV) Runs 1+2



CMS PAS HIG-17-013

Excess here mostly driven by class 1 (&2) at 13 TeV

χ^2 probability for the seven individual values to be compatible with a single signal hypothesis: 41%

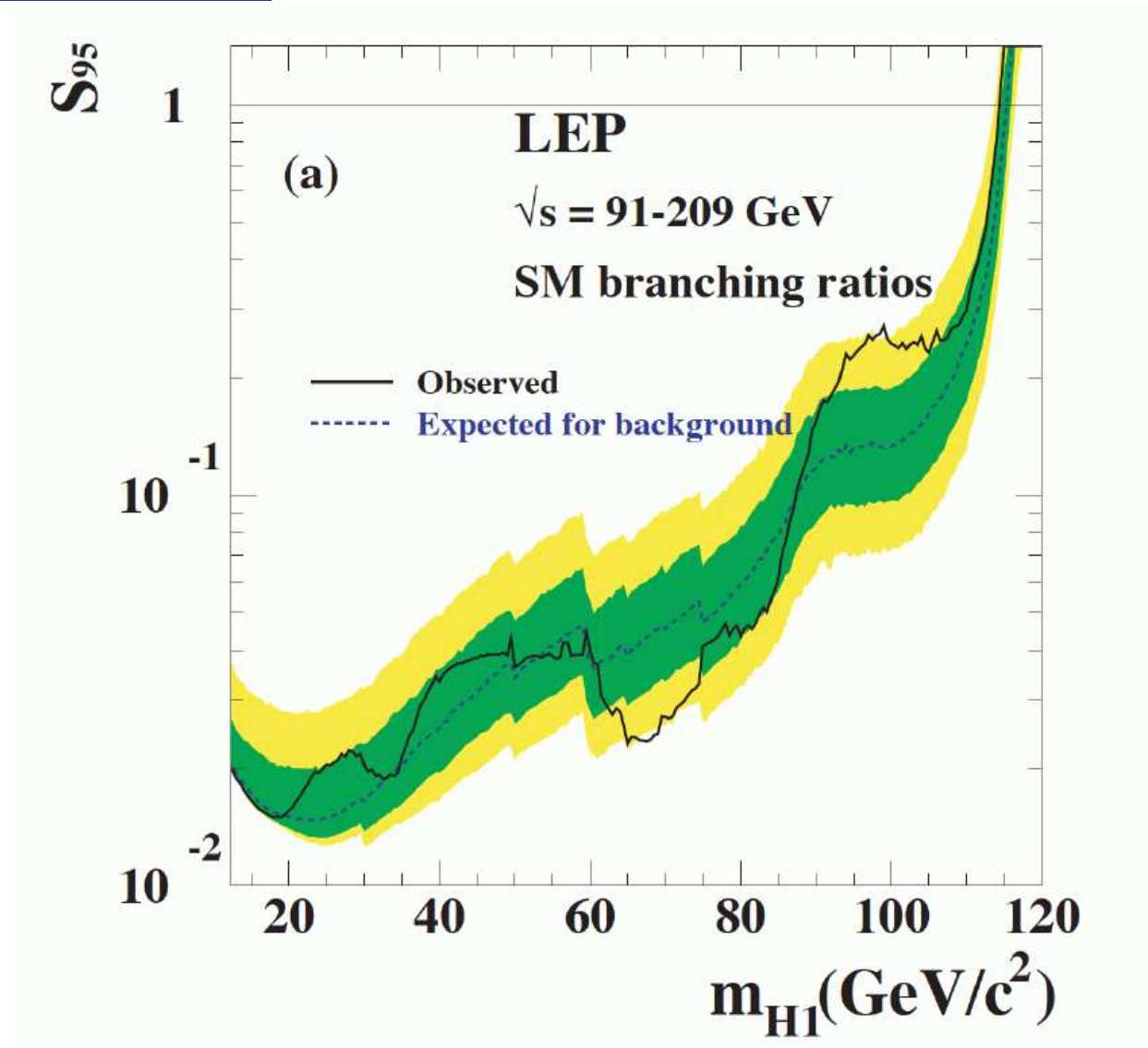
- ‘Signal’ strengths for the 7 event classes and overall, in the 8 TeV+13TeV combination, fixing $m_H=95.3 \text{ GeV}$
- More data are required to ascertain the origin of this excess

S. Gascon-Shotkin HDays17, Santander, ES Sept. 22 2017

55

$$\mu_{\text{CMS}}(96 \text{ GeV}) = [\sigma(pp \rightarrow h_1) \times \text{BR}(h_1 \rightarrow \gamma\gamma)]_{\text{exp/SM}} = 0.6 \pm 0.2$$

What was seen at LEP?



$$\mu_{\text{LEP}}(98 \text{ GeV}) = \left[\sigma(e^+e^- \rightarrow Zh_1) \times \text{BR}(h_1 \rightarrow b\bar{b}) \right]_{\text{exp/SM}} = 0.117 \pm 0.057$$

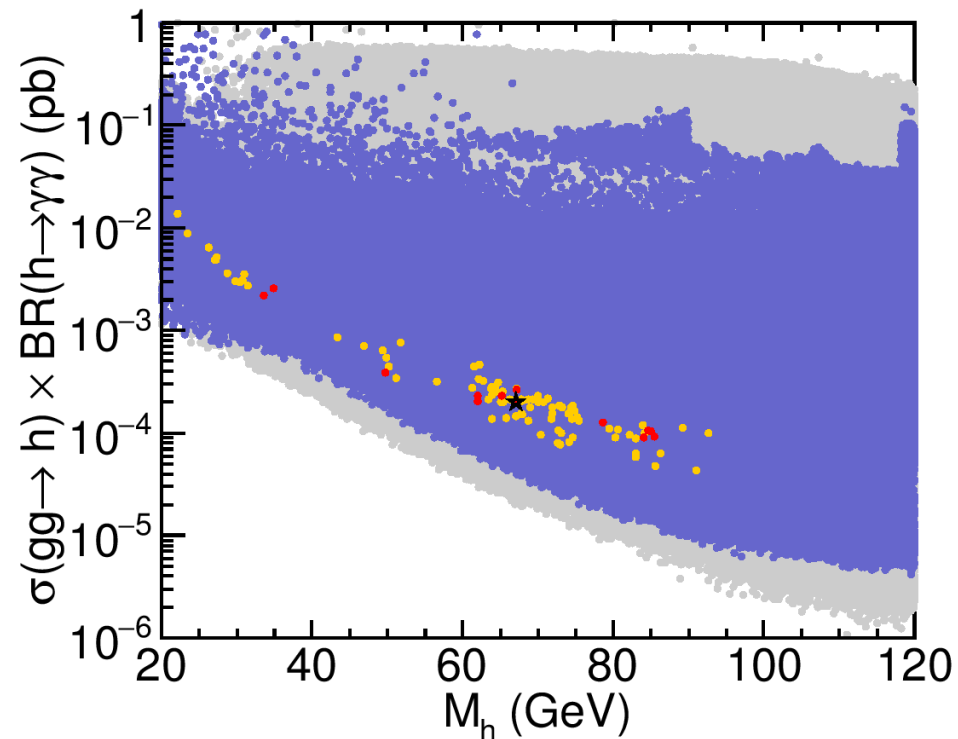
Should we get excited?

⇒ according to CMS no!

⇒ let's wait for ATLAS, ETA summer '18

Which model fits?

[P. Bechtle, H. Haber, S.H., O. Stål, T. Stefaniak, G. Weiglein, L. Zeune '16]



⇒ not the MSSM

⇒ 2HDM? NMSSM?

Check the $\mu\nu$ SSM

$\mu\nu$ SSM: [*D. Lopez-Fogliani, C. Muñoz '06*]

$\mu\nu$ SSM: NMSSM + well motivated RPV (in simple terms)
 \Rightarrow EW scale seesaw to reproduce the neutrino data

Check the $\mu\nu$ SSM

$\mu\nu$ SSM: [D. Lopez-Fogliani, C. Muñoz '06]

$\mu\nu$ SSM: NMSSM + well motivated RPV (in simple terms)
 \Rightarrow EW scale seesaw to reproduce the neutrino data

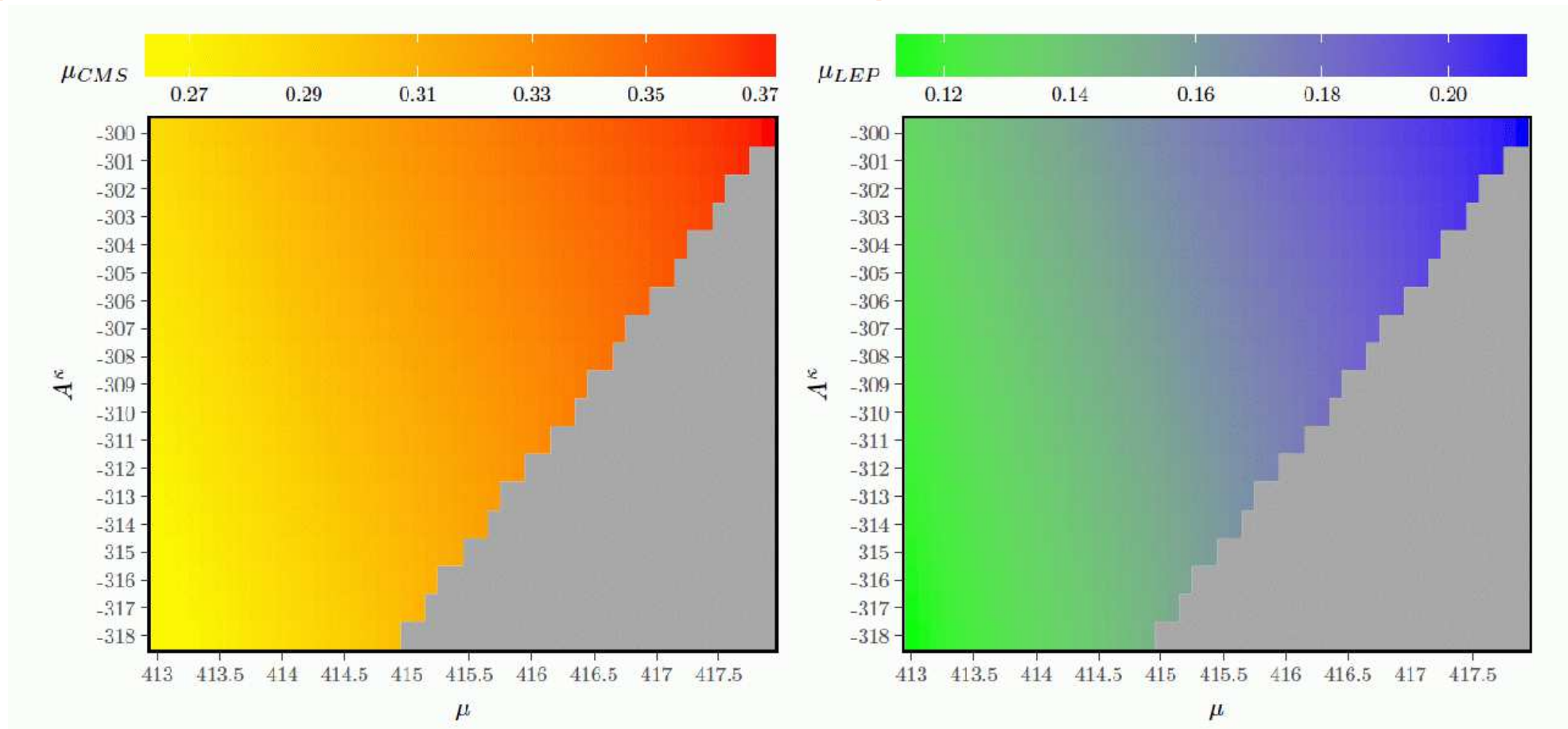
Can the $\mu\nu$ SSM explain the two “excesses”?

[T. Biekötter, S.H., C. Muñoz, arXiv:1712.07475]

v_{iL}	Y_i^ν	A_i^ν	$\tan\beta$	μ	λ	A^λ	κ	A^κ	M_1
$\sqrt{2} \cdot 10^{-5}$	10^{-7}	-1000	2	[413; 418]	0.6	956.035	0.035	[-300; -318]	100
M_2	M_3	$m_{\tilde{Q}_{iL}}^2$	$m_{\tilde{u}_{iR}}^2$	$m_{\tilde{d}_{iR}}^2$	A_1^u	$A_{2,3}^{u,d}$	$(m_e^2)_{ii}$	A_{33}^e	$A_{11,22}^e$
200	1500	800^2	800^2	800^2	0	0	800^2	0	0

Can the $\mu\nu$ SSM explain the two “excesses”?

[*T. Biekötter, S.H., C. Muñoz, arXiv:1712.07475*]



⇒ **YES, WE CAN! :-)**
(at the $1 - 1.5\sigma$ level)

5. Conclusinos

- **SUSY** is (still) the best-motivated BSM scenario
 - constrained models: CMSSM, NUHM1, NUHM2, SU(5), mAMSB
 - general models: pMSSM11, ...
- **MasterCode**: LHC, Higgs, EWPO, BPO, CDM $\Rightarrow \chi^2$ evaluation

Model	Min. χ^2/dof	χ^2 -prob. (p -value)
GUT based models	(30...33)/(18...23)	$\sim 11\%$
pMSSM11	21.0/20	33%

\Rightarrow pMSSM11: model with higher χ^2 -probability
model with interesting (good) ILC (CLIC) prospects
detailed LHC analysis tbd!

- $\phi_{96} \rightarrow \gamma\gamma$: new CMS result possibly interesting!
 $\mu\nu$ SSM can explain CMS and LEP “excesses”

Further Questions?

